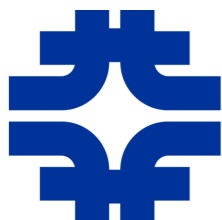


# Beam Cooling (Emittance Reduction)



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- Electron Cooling
- Stochastic Cooling
- Ionization Cooling



# Electron Cooling\*



The appeal of electron cooling is easy to illustrate. In conventional kinetic theory, the gas temperature is related to the mean energy of the molecules by

$$\frac{\langle p^2 \rangle}{2m} = \frac{3}{2}kT.$$

So for an ion beam, one can define a “temperature” for each degree of freedom by

$$\frac{\langle p_{x0}^2 \rangle}{m}, \quad \frac{\langle p_{y0}^2 \rangle}{m}, \quad \frac{\langle p_{z0}^2 \rangle}{m},$$

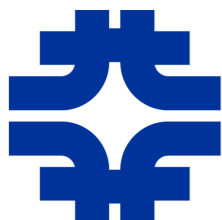
where the Boltzmann constant has been suppressed. The subscript “0” implies that the momenta are measured with respect to the rest frame of the beam centroid.

$$T_x = mc^2 \left( \frac{v}{c} \right)^2 \gamma^2 (\sigma')^2 = mc^2 \left( \frac{v}{c} \right) \gamma \frac{\epsilon_N}{\pi \beta},$$

$$T_s = mc^2 \left( \frac{v}{c} \right)^2 \sigma_p^2, \quad \sigma_p^2 \equiv \left\langle \left( \frac{\Delta p}{p} \right)^2 \right\rangle.$$



LEIR ring, CERN



\*G.1. Budker, *Proc. Intl. Symp. on Electron and Positron Storage Rings*, Saclay, 1966, p. 11-1-1

In electron cooling, an electron beam with very small emittance is made to travel at the same speed as the proton/ion beam that has a relatively high emittance (transverse temperature).

The cool electrons then exchange transverse energy with the hot protons/ions, and then the electrons are discarded (and new electrons re-generated).

## Electron Cooling Time

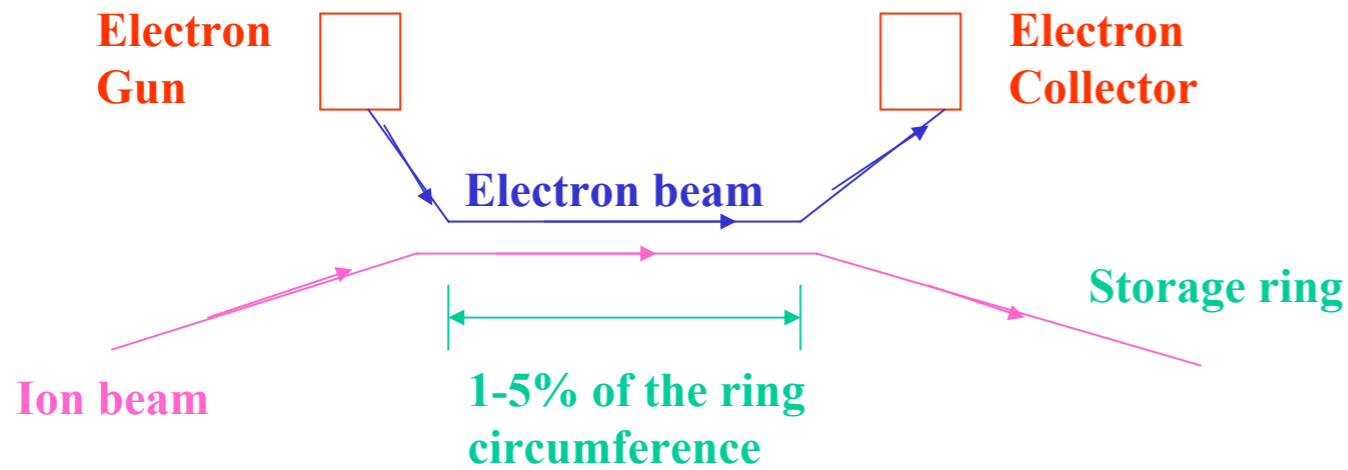
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first estimate:  
(Budker 1967) 
$$\tau = \frac{3}{8\sqrt{2\pi}n_e Q^2 r_e r_i c L_C} \left( \frac{k_B T_e}{m_e c^2} + \frac{k_B T_i}{m_i c^2} \right)^{3/2}$$

for large relative velocities

cooling time 
$$\tau_z \propto \frac{A}{Q^2} \frac{1}{n_e \eta} \beta^3 \gamma^5 \theta_z^3 \begin{cases} \theta_{x,y} = \frac{v_{x,y}}{\gamma \beta c} \\ \theta_{\parallel} = \frac{v_{\parallel}}{\gamma \beta c} \end{cases}$$

# Electron Cooling



To cool  $W = 8$  GeV proton beam requires an electron beam with  $W = 4.3$  MeV must have same speeds:

$$\gamma_p = \gamma_e$$

$$W_p m_e / m_p = W_e$$

## Electron cooling system setup at Fermilab

Pelletron  
(MI-31 building)



Cooling section  
solenoids  
(MI-30 straight section)

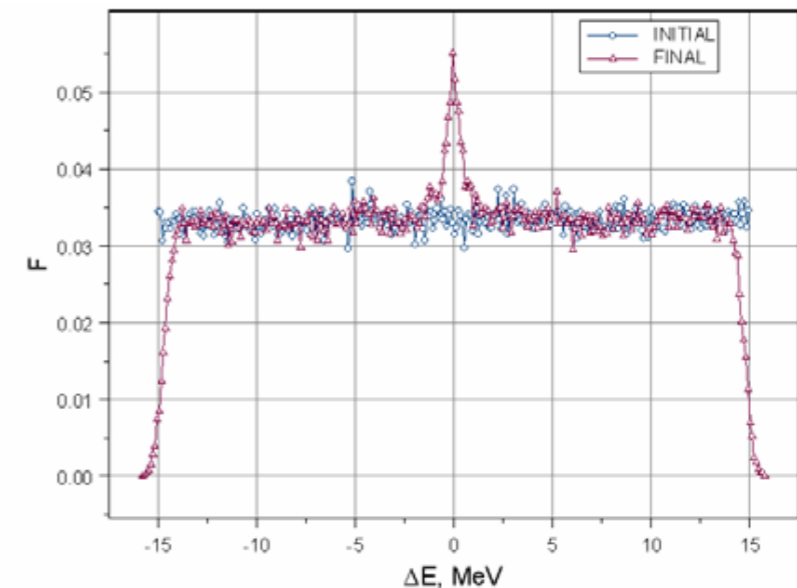


Figure 5: The momentum distribution (arb. units) as a function of antiproton energy deviation (simulation by MOCAC code [5]). The initial distribution is uniform in energy. The final distribution is plotted after 30 minutes.

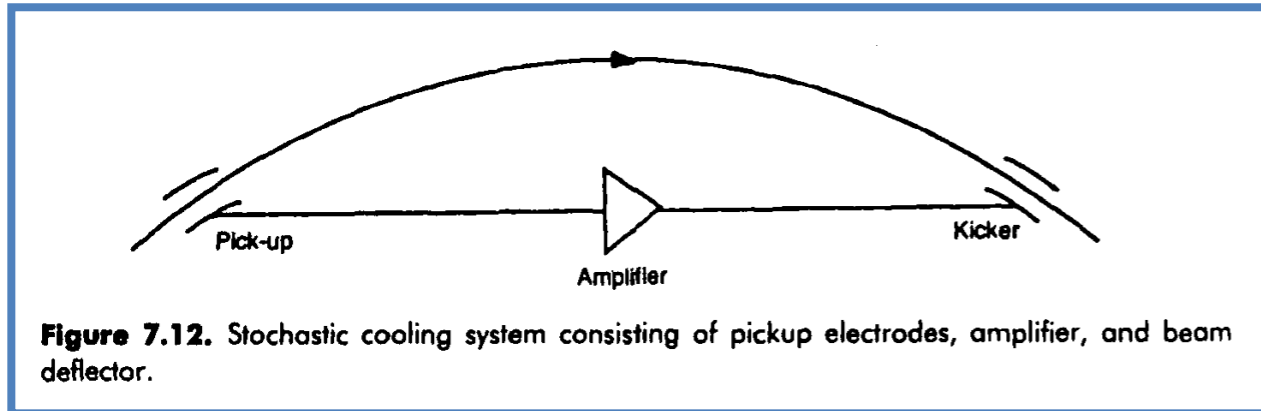


# Transverse Stochastic Cooling\*



$k$  samples in the ring

## Transverse Cooling



$$f_{\max} = \frac{v}{\lambda_{\min}} = \frac{kv}{2C} = \frac{k}{2T}, \quad (7.136)$$

where  $T$  is the revolution period. For a system with a flat frequency response from  $f = 0$  to  $f = W$ ,  $W$  determines  $f_{\max}$ . So the number of particles in a sample, in terms of the bandwidth  $W$ , is given by

$$N_s = \frac{N}{k} = \frac{N}{2TW}. \quad (7.137)$$

where in the second step we have assumed that the various  $x_i$  are uncorrelated. Thus, keeping terms up to first order in  $1/N_s$ , we have for the rate of change of  $\langle x^2 \rangle$

$$\frac{d\langle x^2 \rangle}{dn} = -\frac{2g}{N_s} \langle x^2 \rangle + \frac{g^2}{N_s} \langle x^2 \rangle. \quad (7.143)$$

The cooling rate is then

$$\frac{1}{\epsilon} \frac{d\epsilon}{dn} = -\left( \frac{2g - g^2}{N_s} \right), \quad (7.144)$$

or, in terms of time,

$$\frac{1}{\tau} \equiv -\frac{1}{\epsilon} \frac{d\epsilon}{dt} = -\frac{1}{\epsilon} \frac{d\epsilon}{dn} \frac{1}{T} = \frac{2g - g^2}{N_s T} = \frac{2W}{N} (2g - g^2). \quad (7.145)$$

Averaging over many samples,  $\langle x_n \rangle = 0$  and so

$$\frac{1}{\langle x^2 \rangle} \frac{d\langle x^2 \rangle}{dn} = [-2g + g^2(1 + U)] \frac{1}{N_s}, \quad (7.149)$$

$$M = \frac{T_s}{\Delta T}, \quad (7.150)$$

where  $T_s = (N_s/N)T = 1/(2W)$  is the sample time, and  $\Delta T$  is the change in the revolution period due to the momentum deviation  $\Delta p/p$ . Then

$$M = \frac{1}{2WT|\eta|(\Delta p/p)}. \quad (7.151)$$

For ideal mixing,  $M = 1$ . Intuitively, one would expect the cooling rate to degrade by a factor of  $M$  as we depart from perfect mixing. Actually, this factor of  $M$  appears only in the incoherent term, and so the emittance decreases according to

$$\epsilon = \epsilon_0 e^{-t/\tau}, \quad (7.152)$$

where we have for the cooling rate

$$\frac{1}{\tau} = \frac{2W}{N} [2g - g^2(M + U)].$$

\*D. Mohl, G. Petrucci, L. Thorndahl, and S. van der Meer, "Physics and Technique of Stochastic Cooling," Physics Reports **58**, No.2 (1980).



# Longitudinal Stochastic Cooling

## Cooling

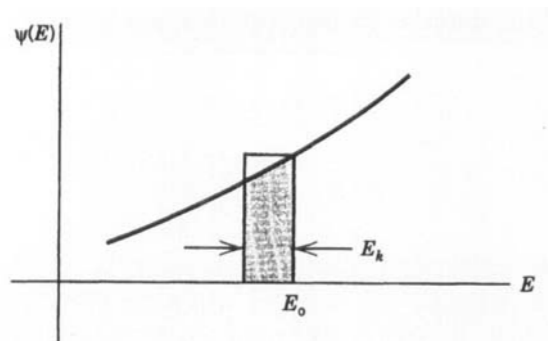


Figure 7.13. Particle density function  $\psi(E)$ .

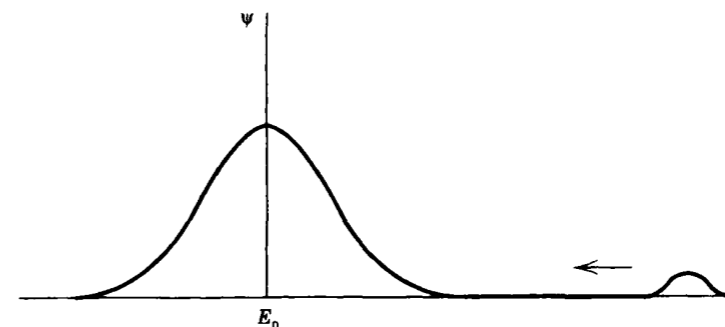


Figure 7.14. Momentum stacking.

As in the discussion of beam-gas scattering, the flux arising from diffusion can be written in the form

$$\vec{J} = -D \nabla \psi, \quad (7.154)$$

where  $\vec{J}$  is the particle flux. In the case under consideration here, since energy is the only degree of freedom,

$$J = -D(E) \frac{\partial \psi}{\partial E}, \quad (7.155)$$

where the diffusion “constant” may be a function of energy. To this, we must add coherent forces. If the rate of energy gain is  $C(E)$ , then we must add  $\psi C(E)$  to the flux, obtaining

$$J = C(E)\psi - D(E) \frac{\partial \psi}{\partial E}. \quad (7.156)$$

in which the flux is zero. Suppose there is a coherent force driving particles toward some central energy  $E_0$ , where the force is proportional to the energy deviation  $E - E_0$ , and suppose that the diffusion force is a constant. Then  $C(E) = -\alpha(E - E_0)$  and  $D(E) = D_0$ . So this static situation is described by

$$J = -\alpha(E - E_0)\psi - D_0 \frac{\partial \psi}{\partial E} = 0. \quad (7.163)$$

The solution to the above equation is the Gaussian

$$\psi = \psi_0 e^{-\alpha(E - E_0)^2 / 2D_0}. \quad (7.164)$$

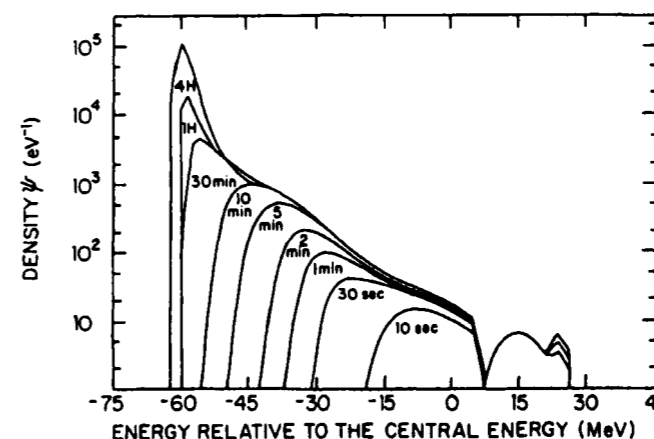


Figure 7.15. Design curves for antiproton energy density at FNAL Accumulator Ring, showing development of the “ $\bar{p}$  stack” over time. From Tollestrup and Dugan, with permission.

So we have

$$\frac{\partial \psi}{\partial E} = -\frac{e^2 \psi}{4J_0 T^2 A} + \frac{e^2 \psi}{2J_0 T^2 A} = \frac{e^2 \psi}{4AT^2 J_0} \equiv \frac{\psi}{E_d}, \quad (7.171)$$

or

$$\psi = \psi_0 e^{(E - E_0)/E_d}. \quad (7.172)$$

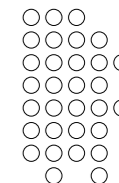


# Ionization Cooling\*



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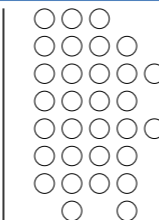
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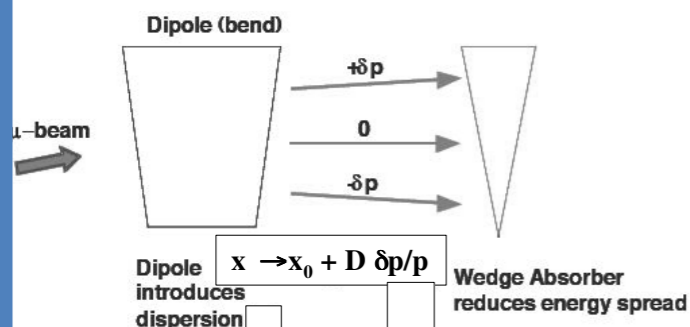
## Applications of Ionization Cooling

### Ionization Cooling

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increased longitudinal cooling  
by longitudinal-transverse emittance exchange



$$\frac{d\sigma_E^2}{ds} = -2 \frac{\partial(dE/ds)}{\partial E} \sigma_E^2 + \frac{d\langle \Delta E_{rms}^2 \rangle}{ds}$$

cooling term                  heating term

cooling, if  $\frac{\partial(dE/ds)}{\partial E} > 0$

#### emittance exchange

increased longitudinal cooling

$$\frac{\partial \frac{dE}{ds}}{\partial E} \Rightarrow \frac{\partial \frac{dE}{ds}}{\partial E} \Big|_0 + \frac{dE}{ds} \frac{D\rho'}{\beta c p \rho_0}$$

reduced transverse cooling

$$\frac{d\epsilon_N}{ds} = -\frac{1}{\beta^2 E} \frac{dE}{ds} \left(1 - \frac{D\rho'}{\rho_0}\right) \epsilon_N$$

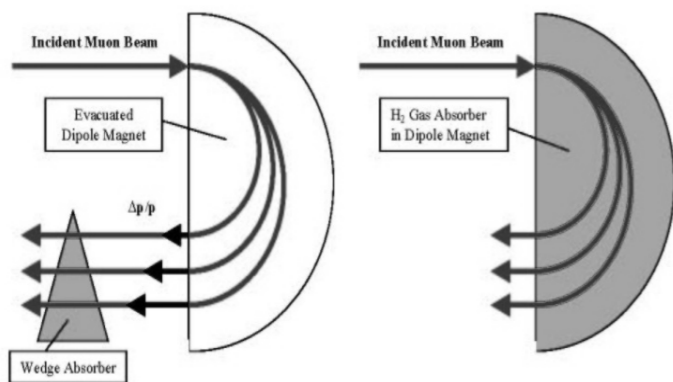
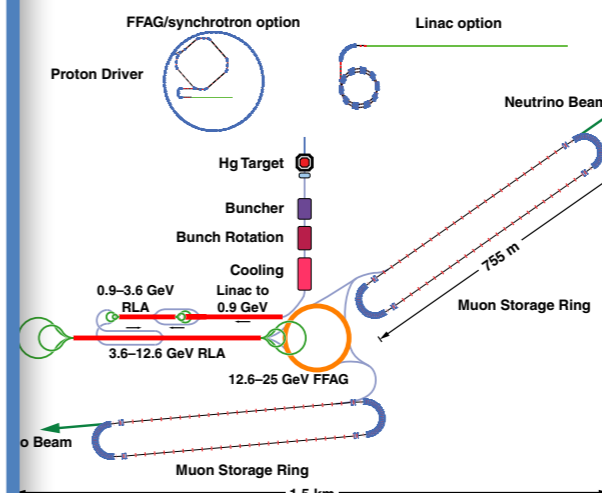


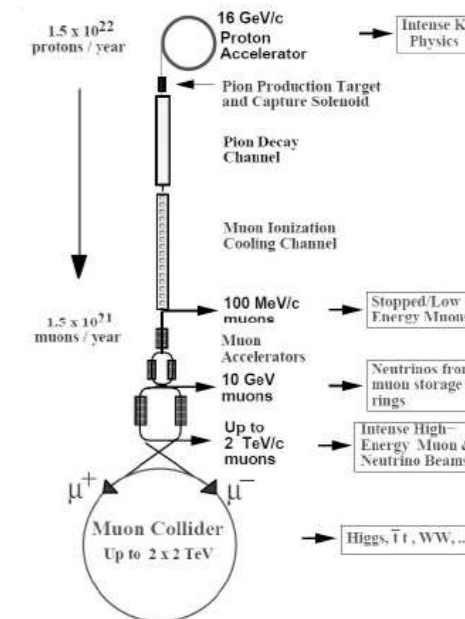
Figure 1. Use of a Wedge Absorber for Emittance Exchange

Figure 2. Use of Continuous Gaseous Absorber for Emittance Exchange

### Neutrino Factory

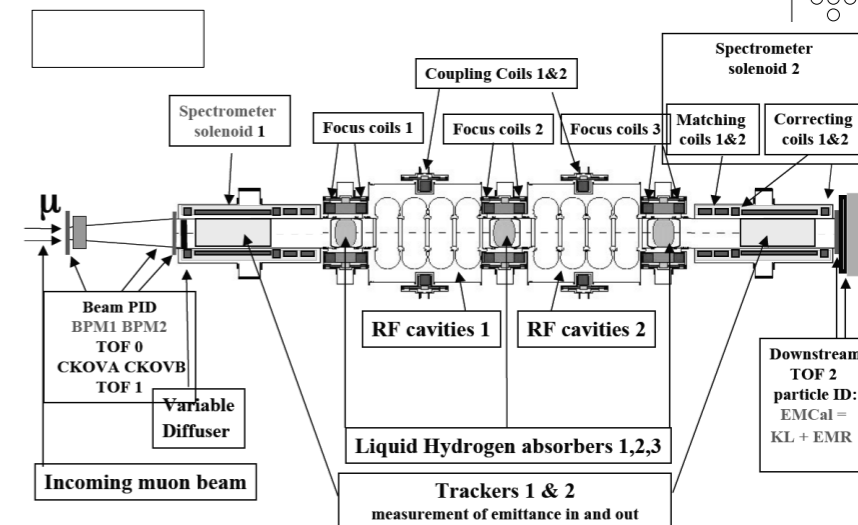


### Muon Collider

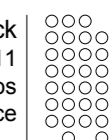


### MICE

Muon Ionization Cooling Experiment at RAL



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\* G.I. Budker, in: Proceedings of 15th International Conference on High Energy Physics, Kiev, 1970