

Field of a moving particle in a dielectric

- The E- and B-field associated to a uniformly moving particle in vacuum are given by

$$\vec{E}(\vec{x}, \omega) = \sqrt{\frac{2}{\pi}} \frac{q}{v} \left[\frac{\lambda}{\epsilon} K_1(b\lambda) \hat{x} - i \frac{\omega}{v} (\epsilon\beta - 1) K_0(b\lambda) \hat{z} \right].$$

- and

$$\vec{B} = \sqrt{\frac{2}{\pi}} \frac{q}{c} \lambda K_1(b\lambda) \hat{y}.$$

- where $\lambda \equiv \frac{\omega}{v} \sqrt{1 - \epsilon \frac{v^2}{c^2}}$
- Previously we considered the case $b\lambda \ll 1$



Field in the limit $b\lambda \gg 1$

- We now consider the extreme case $b\lambda \gg 1$. The modified Bessel functions have the asymptotic expansion:

$$K_0(b\lambda) = K_1(b\lambda) = \sqrt{\frac{\pi}{2}} \frac{e^{-b\lambda}}{\sqrt{b\lambda}}$$

- And the field associated to a uniformly moving charged particle takes the form

$$\begin{aligned}\vec{E}(\vec{x}, \omega) &= \frac{q}{v} \frac{e^{-b\lambda}}{\sqrt{b\lambda}} \left(\frac{\lambda}{\epsilon} \hat{x} - i \frac{\omega}{v} \left(\frac{1}{\epsilon} - \beta^2 \right) \hat{z} \right), \\ \vec{B}(\vec{x}, \omega) &= \frac{q}{v} \frac{e^{-b\lambda}}{\sqrt{b\lambda}} \hat{y}.\end{aligned}$$



Cerenkov condition

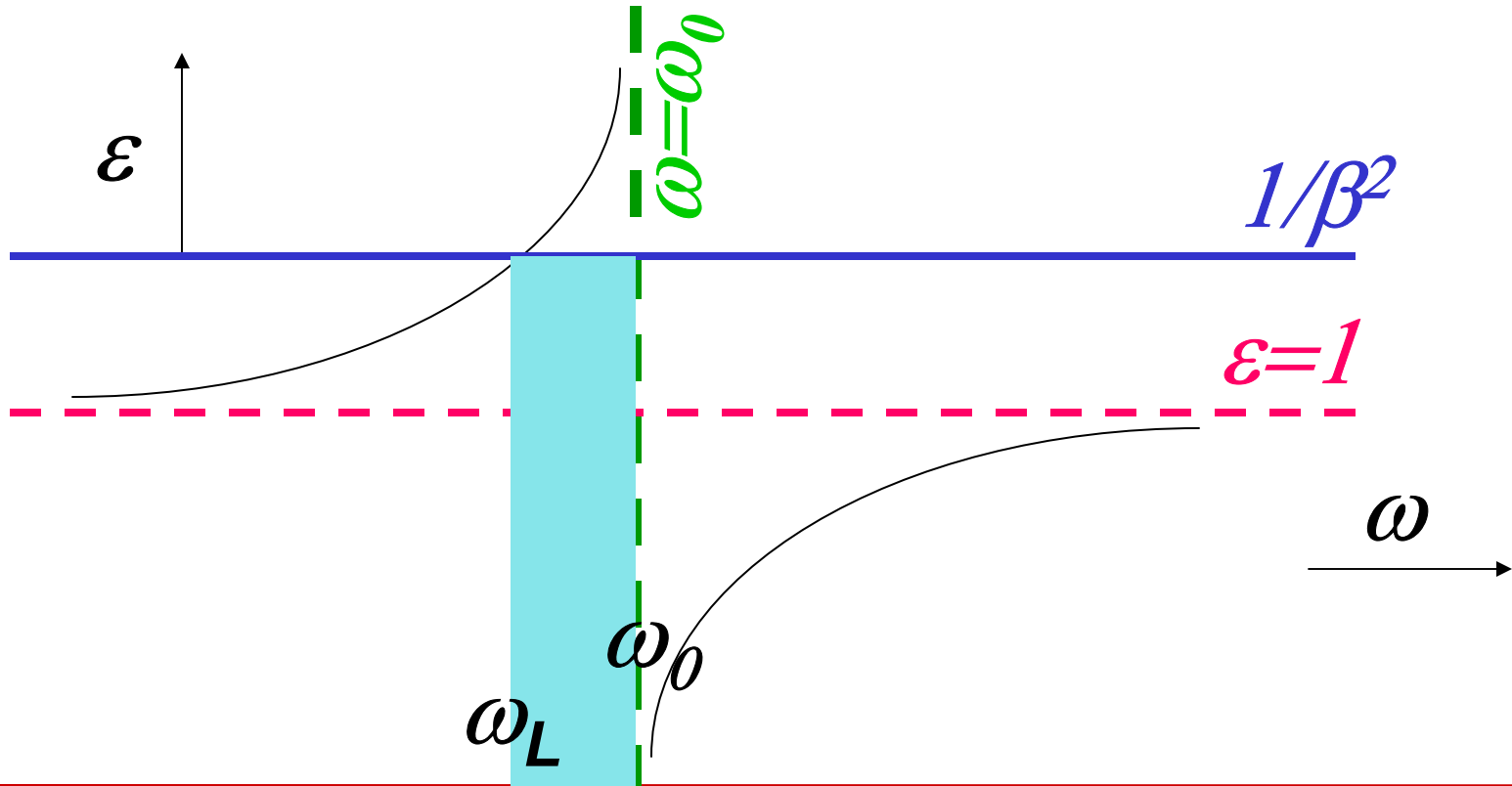
- We saw that to get radiation (or energy loss) we needed either λ or ϵ to be complex.
- When investigating dielectric screening effects we considered the case $\epsilon \in \mathbb{C}$
- We now consider the case where λ is a pure imaginary number and $\epsilon \in \mathbb{R}$.
- from $\lambda = \frac{\omega}{v} \sqrt{1 - \epsilon(\omega)\beta^2}$
- $\lambda \in \mathbb{I}, 1 - \epsilon\beta^2 < 0 \Rightarrow \epsilon\beta^2 > 1$, this is the Cerenkov condition.



Cerenkov condition

- Consider the model for permittivity

$$\epsilon(\omega) = 1 + \frac{\omega_p^2}{\omega_0^2 - \omega^2 - i\omega\Gamma}$$



Energy loss

- Explicit the asymptotic form of the e.m. field in the energy loss equation.

$$\begin{aligned}\frac{d\mathcal{E}_f}{dz} &= \mathcal{R}e \left\{ \int_0^{+\infty} d\omega \frac{2}{\pi} \frac{q^2}{v^2} [i\omega(1/\epsilon - \beta^2)\lambda^*b] K_0(\lambda b) K_1(\lambda^*b) \right\} \\ &= \frac{2}{\pi} \frac{q^2}{v^2} \mathcal{R}e \left\{ \int_0^{+\infty} d\omega (i\omega\lambda^*b)(1/\epsilon - \beta^2) K_0(\lambda b) K_1(\lambda^*b) \right\}\end{aligned}$$

- gives

$$\begin{aligned}\frac{d\mathcal{E}_f}{dz} &= \frac{q^2}{v^2} \mathcal{R}e \left(\int_0^\infty d\omega (i\omega\lambda^*b) \left(\frac{1}{\epsilon} - \beta^2 \right) K_0(b\lambda) K_1(b\lambda^*) \right) \\ &= \frac{q^2}{v^2} \mathcal{R}e \left(\int_0^\infty d\omega (i\omega\lambda^*b) \left(\frac{1}{\epsilon} - \beta^2 \right) \frac{e^{-b(\lambda+\lambda^*)}}{b\sqrt{\lambda\lambda^*}} \right) \\ &= \frac{q^2}{v^2} \mathcal{R}e \left(\int_0^\infty i\omega \sqrt{\frac{\lambda^*}{\lambda}} \left(\frac{1}{\epsilon} - \beta^2 \right) \right)\end{aligned}$$



Frank-Tamm energy loss formula

- But $\lambda \in I$ so $\lambda^*/\lambda = -1$ so finally

$$\frac{d\mathcal{E}_f}{dz} = \frac{q^2}{v^2} \int_{\omega_L}^{\omega_0} d\omega \omega \left(1 - \frac{1}{\epsilon \beta^2} \right)$$

- This is Frank-Tamm formula derived in 1937.
- History:
 - Cerenkov observed the radiation in Vavilov's labs (1934)
 - Frank and Tamm explained the effect (1937)
 - Cerenkov, Frank and Tamm share Nobel prize (1958)



Direction of propagation

- The direction of the wave is given by \mathbf{k} , \mathbf{k} perpendicular to \mathbf{E} and \mathbf{B} .
Let θ_c be the angle between the velocity of the particle and \mathbf{k} then

$$\begin{aligned}\cos \theta_c &= \frac{|E_x|}{|E|} = \frac{E_x}{\sqrt{E_x^2 + E_z^2}} \\ &= \frac{\frac{\lambda}{\epsilon}}{\left[\left(\frac{\lambda}{\epsilon} \right)^2 - \frac{\omega^2}{v^2} \left(\frac{1}{\epsilon} - \beta^2 \right)^2 \right]^{1/2}}\end{aligned}$$

- From $\lambda^2 = (\omega/v)^2(1 - \epsilon\beta^2)$

$$\cos \theta_c = \frac{1}{\sqrt{1 - 1 + \beta^2 \epsilon}} = \frac{1}{\beta \sqrt{\epsilon}} = \frac{c_m}{v}$$

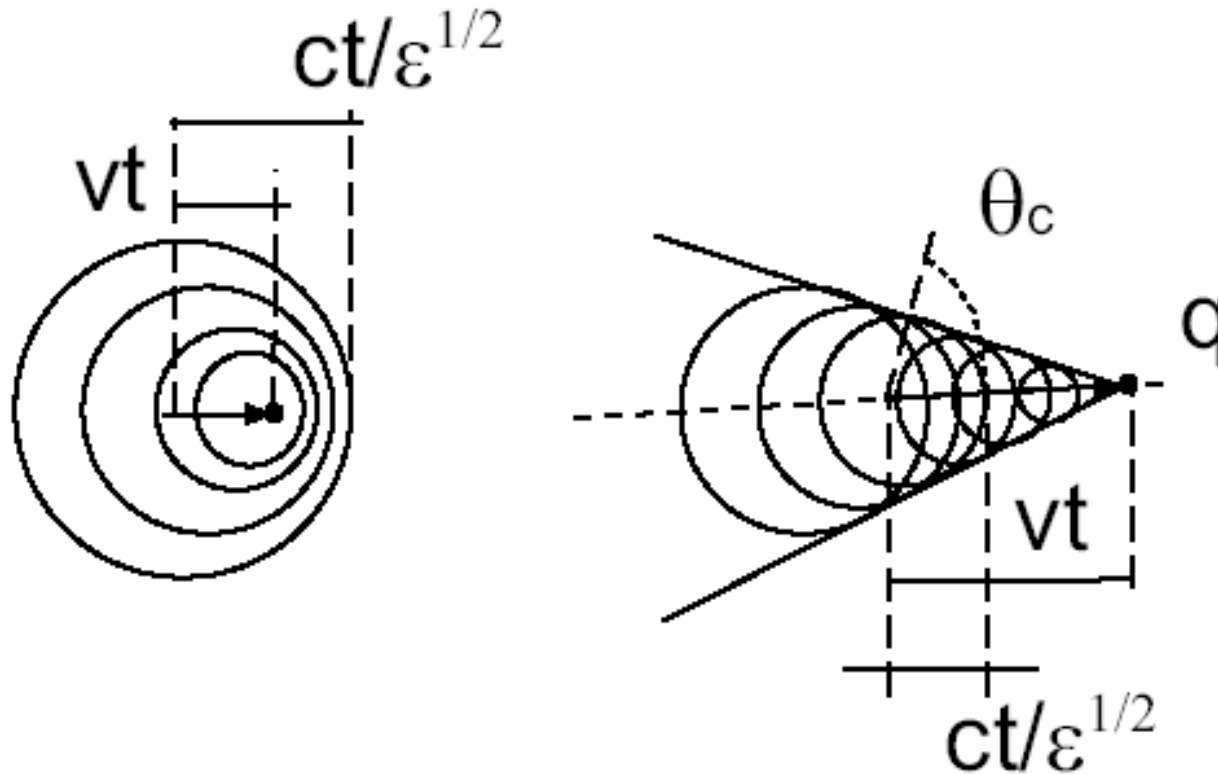
*Velocity of light in
the considered medium*

$c_m < c$ so $\cos \theta_c < 1$ and $\theta \in \mathbb{R}$.



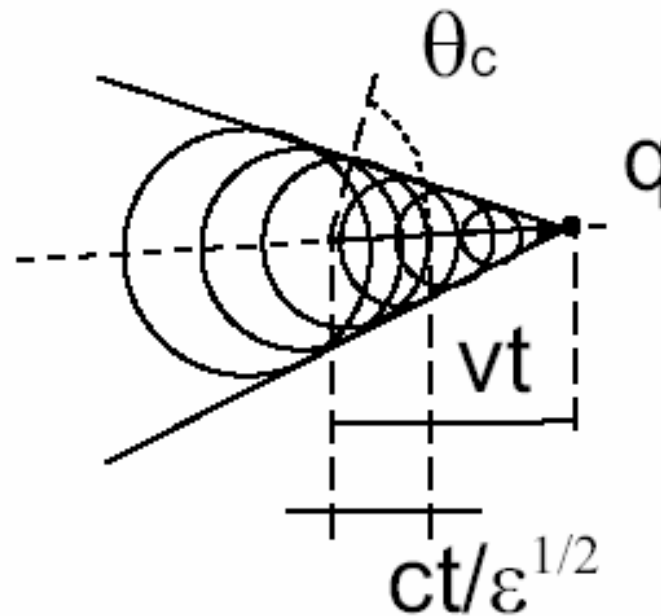
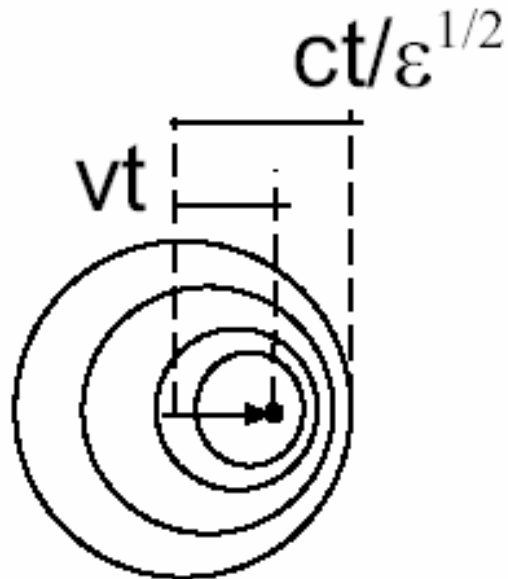
Shock wave feature

- Cerenkov radiation consists of a shock wave
- Effect similar to the Mach effect



Shock wave feature I

- Cerenkov radiation consists of a shock wave
- Effect similar to the Mach effect



Shock wave feature II

- The Shock wave feature inferred geometrically can be derived from the wave equation

$$\begin{aligned}\left(k^2 - \frac{\omega^2}{c_m^2}\right) \sqrt{\epsilon} \Phi(\vec{k}, \omega) &= \frac{4\pi}{\sqrt{\epsilon}} \rho(\vec{k}, \omega) \\ \left(k^2 - \frac{\omega^2}{c_m^2}\right) \sqrt{\epsilon} \vec{A}(\vec{k}, \omega) &= \frac{4\pi}{c_m} \vec{J}(\vec{k}, \omega)\end{aligned}$$

- So A takes the same form as in vacuum under “renormalization”

$$q \rightarrow q/\sqrt{\epsilon}, \quad c \rightarrow c_m.$$

- So we can directly write the potentials as

$$\begin{pmatrix} \sqrt{\epsilon} \Phi(\vec{x}, t) \\ \vec{A}(\vec{x}, t) \end{pmatrix} = \frac{q}{\sqrt{\epsilon}} \frac{1}{[\kappa R]_{ret}} \begin{pmatrix} 1 \\ \frac{\vec{v}}{c_m} \end{pmatrix}$$



Shock wave feature III

- Consider

$$\left. \begin{aligned} \vec{\zeta} &= \vec{x} - \vec{v}t, \\ \vec{R} &= \vec{x} - \vec{x}(t') = \vec{x} - \vec{v}t' \end{aligned} \right\} \vec{R} = \vec{x} - \vec{v}t + \vec{v}(t - t') = \vec{\zeta} + \vec{v}(t - t').$$

- On another hand $\underbrace{t - t'}_{\text{causality}} = \frac{R(t')}{c_m} = \frac{|\vec{\zeta} + \vec{v}(t - t')|}{c_m}.$

- So $(t - t')^2 = \frac{1}{c_m^2} [\zeta^2 + 2 \vec{\zeta} \cdot \vec{v}(t - t') + v^2(t - t')^2].$
 $\Rightarrow (v^2 - c_m^2)(t - t')^2 + 2 \vec{\zeta} \cdot \vec{v}(t - t') + \zeta^2 = 0;$

- Solve for $(t - t')$:

$$(t - t')_{\pm} = \frac{-\vec{\zeta} \cdot \vec{v} \pm \sqrt{(\vec{\zeta} \cdot \vec{v})^2 - (v^2 - c_m^2)\zeta^2}}{v^2 - c_m^2}.$$



Shock wave feature IV

$$(t - t')_{\pm} = \frac{-\vec{\zeta} \cdot \vec{v} \pm \sqrt{(\vec{\zeta} \cdot \vec{v})^2 - (v^2 - c_m^2)\zeta^2}}{v^2 - c_m^2}.$$

- For Cerenkov radiation ($v > c_m$), to obtain $t-t'$ real positive we need $\zeta \cdot v > 0$ and $(\zeta \cdot v)^2 > (v^2 - c_m^2)\zeta^2$.
- That is $\zeta v \cos \theta < 0$ or $\theta > \pi/2$, and $\cos^2 \theta > 1 - c_m^2/v^2$.
- So $\theta > \arccos(-\sqrt{1 - (c_m/v)^2})$,
- Thus potentials and fields exist at time t only within a cone with apex lying at the present time position of the particle and apex angle

$$\pi - \arccos(-\sqrt{1 - (c_m/v)^2})$$



Shock wave feature V

- The 4-potential is given by $A^\alpha = A_-^\alpha + A_+^\alpha$
- The denominator is

$$\begin{aligned}
 [\kappa R]_{ret} &= |(1 - 1/c_m \vec{v} \cdot \hat{n}) \vec{R}| = |\vec{R} - \frac{\hat{n}}{c_m} \vec{v} \cdot [\vec{\zeta} + \vec{v}(t - t')]| \\
 &= |\vec{R} - \frac{\hat{n}}{c_m} \vec{\zeta} \cdot \vec{v} - \frac{\hat{n}}{c_m} v^2 (t - t')| \\
 &= |\hat{n} [c_m (t - t') - \frac{\vec{\zeta} \cdot \vec{v}}{c_m} - \frac{v^2}{c_m} (t - t')]| \\
 &= \frac{1}{c_m} |(c_m^2 - v^2)(t - t') - \vec{\zeta} \cdot \vec{v}|.
 \end{aligned}$$

- Expliciting (t-t') we gives

$$[\kappa R]_{ret} = \frac{\zeta}{c_m} \sqrt{c_m^2 - v^2 \sin^2 \theta} = \zeta \sqrt{1 - \frac{v^2}{c_m^2} \sin^2 \theta}.$$



Shock wave feature VI

- So finally the potentials are

$$\left(\begin{array}{c} \sqrt{\epsilon}\Phi(\vec{x}, t) \\ \vec{A}(\vec{x}, t) \end{array} \right) = \frac{2q}{\sqrt{\epsilon}} \frac{1}{\zeta \sqrt{1 - \frac{v^2}{c_m^2} \sin^2 \theta}} \left(\begin{array}{c} 1 \\ \frac{\vec{v}}{c_m} \end{array} \right)$$

singularity

- In practice the singularity is smeared by the frequency-dependence of $\epsilon(\omega)$ which implies $c_m(\omega)$...

