Before we do anything else

## Let's define some units

SI Units

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\begin{abstract}

\begin{abstract}


#### Abstract




\end{abstract}

\end{abstract}

$\qquad$
$\qquad$


## SI Units

Hydrogen Atom


## Electron mass = $9 \times 10^{-31} \mathrm{~kg}$

## Units <br> Jnits



## 



## 

## 



## 



.

## Units

Hydrogen Atom



## This takes 0.4 seconds

## Units



Neutral pion
lifetime =
$8 \times 10^{-17}$ s

## Do not use convenient particle physics units!

## Smallest things in the Universe

by u/xvmir

## Electron <br>  <br> Quark <br> $X$ on the mobile ad

## Other issues with SI units

$$
\frac{-\hbar^{2}}{2 m} \frac{\partial^{2} u(x)}{\partial x^{2}}+V(x) u(x)=E u(x)
$$

$\hbar$ are pesky and annoying!

$$
\gamma=\frac{1}{\sqrt{1-\frac{v^{2}}{c^{2}}}}
$$

$$
E^{2}=(p c)^{2}+\left(m c^{2}\right)^{2}
$$

so are factors of c everywhere!

## The units that we will use

$$
\hbar=c=1 \quad \begin{gathered}
c=3 \times 10^{8} \mathrm{~m} / \mathrm{s} \\
\hbar=10^{-34} \mathrm{~J} \mathrm{~s}
\end{gathered}
$$

Natural units. Simplifies a lot of notation!

$$
\begin{gathered}
1 \mathrm{GeV}=10^{9} \mathrm{ev}=1.6 \times 10^{-10} \mathrm{~J} \\
\hbar c=0.2 \times 10^{-15} \mathrm{~m}=0.2 \times 10^{-15} \mathrm{fm}
\end{gathered}
$$

|  | Not our choice | In our choice of natural units |
| :--- | :---: | :---: |
| Energy | GeV | GeV |
| Momentum | $\mathrm{GeV} / \mathrm{c}$ | GeV |
| Mass | $\mathrm{GeV} / \mathrm{c}^{2}$ | GeV |
| Time | $\mathrm{hbar} / \mathrm{GeV}$ | $\mathrm{GeV}-1$ |
| Length | $\mathrm{C}^{*} \mathrm{hbar} / \mathrm{GeV}$ | GeV -1 |
| Area | $\left(c^{*} h b a r / G e V\right)^{2}$ | $\mathrm{GeV}^{-2}$ |

## Griffiths disagrees :)



JJ Thompson, who measured the charge to mass ratio of "cathode rays" (aka electrons). Any ideas how he found their velocity and, more importantly, q/m? (1.1 in Griffiths)

## Let's take a step back in time, now (1897)



## Undeflected particle: $q E=q v B, v=(E / B)$

Just magnetic field: $q v B=m v^{2} / R$
$(\mathrm{q} / \mathrm{m})=\mathrm{v} /(\mathrm{BR})=\mathrm{E} /\left(\mathrm{B}^{2} \mathrm{R}\right)$

## On to Rutherford (1911)


"It was quite the most incredible event that has ever happened to me in my life. It was almost as incredible as if you fired a 15 -inch shell at a piece of tissue paper and it came back and hit you." - let to the idea that positive charge in atom must be concentrated in a small space


Quantization was "a purely formal assumption and I really did not give it much thought except that no matter what the cost, I must bring about a positive end." (Turns out that Planck didn't quite realize what he was doing)

## Einstein and the photoelectric effect (1905)



A radical idea! The energy of a photon is quantized, and depends on its frequency (color). Higher intensity light will knock out more electrons, but always of the same energy! Isn't light a wave and not a particle?

## Conservation of energy

$$
E+m=E^{\prime}+\sqrt{p^{2}+m^{2}}
$$

E'



Momentum conservation in x :

$$
E=E^{\prime} \cos \theta+p \cos \phi
$$

Momentum conservation in y:
$0=E^{\prime} \sin \theta-p \sin \phi$

## Momentum conservation in y :



## Momentum conservation in $x$ :

E'

E

$$
\begin{gathered}
E=E^{\prime} \cos \theta+p \cos \phi \\
E=E^{\prime} \cos \theta+p \sqrt{1-\frac{E^{\prime 2}}{p^{2}} \sin ^{2} \theta} \\
\left(E-E^{\prime} \cos \theta\right)^{2}=p^{2}\left(1-\frac{E^{\prime 2}}{p^{2}} \sin ^{2} \theta\right)
\end{gathered}
$$

$$
E^{2}+E^{\prime 2} \cos ^{2} \theta-2 E E^{\prime} \cos \theta=p^{2}-E^{\prime 2} \sin ^{2} \theta
$$

$$
p^{2}=E^{2}+E^{\prime 2}-2 E E^{\prime} \cos \theta
$$

## Conservation of energy again

$$
\begin{gathered}
E+m=E^{\prime}+\sqrt{p^{2}+m^{2}} \\
\left(E+m-E^{\prime}\right)=\sqrt{p^{2}+m^{2}} \quad \text { Plug in } \mathrm{p}^{2} \\
\left(E+m-E^{\prime}\right)^{2}=E^{2}+E^{\prime 2}-2 E E^{\prime} \cos \theta+m^{2} \\
E^{2}+E^{\prime 2}+m^{2}+2 m E-2 m E^{\prime}-2 E E^{\prime}=E^{2}+E^{\prime 2}-2 E E^{\prime} \cos \theta+m^{2} \\
2 m E-2 m E^{\prime}=2 E E^{\prime}(1-\cos \theta) \\
m\left(\frac{1}{\lambda}-\frac{1}{\lambda^{\prime}}\right)=\frac{1}{\lambda \lambda^{\prime}}(1-\cos \theta) \\
\text { Nice to have these units }
\end{gathered}, \begin{gathered}
\\
E^{\prime}=\frac{1}{\lambda} \\
\lambda^{\prime}
\end{gathered}
$$

Finishing up

$$
\begin{aligned}
& \gamma\left(\lambda^{\prime}\right) \wedge^{m}\left(\frac{1}{\lambda}-\frac{1}{\lambda^{\prime}}\right)=\frac{1}{\lambda \lambda^{\prime}}(1-\cos \theta) \\
& \begin{array}{l}
m\left(\lambda^{\prime}-\lambda\right)=(1-\cos \theta) \\
\cdots \lambda^{\prime}=\lambda+\frac{1-\cos \theta}{m}
\end{array} \\
& \text { What happens for different } \\
& \text { angles? Different m? }
\end{aligned}
$$

Note that this assumes particles of light, aka photons! No discussion of waves or interference


Photon as a particle (classical). But what we really mean is that the field is quantized (here providing an attractive force), and the quantized unit of the field transmits some momentum from one object to another

## Early particle detectors - cloud and bubble chambers

A cloud chamber is a collection of
supersaturated vapor of alcohol or water. A charge particle can ionize the vapor; the subsequent ions act as seeds for condensation. Magnetic fields can give the charge and momenta of objects. Need to literally take pictures of the chambers!

Bubble chamber is similar, except it uses superheated liquid instead

## What's

 happening?
## Early particle detectors - cloud and bubble chambers

Glaser invented the bubble chamber (and won a Nobel prize for it).
"Legend has it that while he was on the faculty of the University of Michigan, Glaser was chilling with colleagues over a cold beer, observed the stream of bubbles in his glass, and was inspired to build a device that could track subatomic particles with bubbles. Glaser himself later refuted this story; beer was not his inspiration, although he did use it as a liquid in early prototypes."

## FNAL bubble chamber photo



## What's going on here?

## http://www.symmetrymagazine.org/article/ january-2015/how-to-build-your-own-particledetector



How to build your own particle detector

The scale of the detectors at the Large Hadron Collider is
aimost incomprehensible: They weigh thousands of tons. almost incomprehensible: They weigh thousands of tons, program for an international community of thousands of scientists.
fut particle detectors aren't always so comple some particle detectors are so simple that you can make (and perate) them in your own home.

The Continuously Sensitive Diffusion Cloud Chamber is one such detector. Originally developed at UC Berkeley in 1938 , this ype of detector uses evaporated alcohol to make a 'clowd' that extremely sensitive to passing particles.

Cosmic rays are particles that are constantly crashing into the Earth from space. When they hit Earth's atmosphere, they release a shower of less massive particles, many of which

If anyone knows any enterprising undergrads looking for work, we tried building one a few years ago, but didn't really spend the time or effort. This would be a lot of fun


Strong meson
(intermediate mass) must be the nuclear force carrier with mass $\sim 150 \mathrm{MeV}$. See Griffiths HW 1.2 for why this was only a vague estimate - any ideas how he got it (if you haven't read the textbook?)

## There are two such mesons observed in cosmic rays:

$$
\pi, \mu
$$

VERY different interactions with atomic nuclei (one feels the strong nuclear force, the other does not)



Fig. 1.3 One of Powell's earliest pictures showing the track of a pion in a photographic emulsion exposed to cosmic rays at high altitude. The pion (entering from the left) decays into a muon and a neutrino (the latter is electrically neutral, and leaves no

## From Griffiths

track). (Source: Powell, C. F., Fowler, P. H. and Perkins, D. H. (1959) The Study of Elementary Particles by the Photographic Method Pergamon, New York. First published in (1947) Nature 159, 694.)

## From Griffiths

## Hypothesized by Dirac (we'll see why in a few chapters)



Fig. 1.4 The positron. In 1932. Anderson took this photograph of the track left in a cloud chamber by a cosmic ray particle. The chamber was placed in a magnetic field (pointing into the page), which caused the particle to travel in a curve. But was it a negative charge traveling downward or a positive charge traveling upward? In order to distinguish. Anderson had placed a lead plate across the center of the chamber (the thick horizontal line in the photograph). A

particle passing through the plate slows down, and subsequently moves in a tighter circle. By inspection of the curves, it is clear that this particle traveled upward, and hence must have been positively charged. From the curvature of the track and from its texture, Anderson was able to show that the mass of the particle was close to that of the electron. (Photo courtesy California Institute of Technology.)

Protons vs anti-proton

$$
\begin{aligned}
& p=(u u d) \\
& \bar{p}=(\overline{u u} \bar{d})
\end{aligned}
$$

Neutron vs antineutron (note: neutron has no net electric charge!)

$$
n=(u d d)
$$

$$
\begin{gathered}
e^{+} \operatorname{VS} e^{-} \\
\gamma=\bar{\gamma}
\end{gathered}
$$

$$
\begin{gathered}
A+B \rightarrow C+D \rightarrow \quad \begin{array}{l}
\text { If this then } \\
\text { also... }
\end{array} \\
A \rightarrow \bar{B}+C+D \\
\begin{array}{c} 
\\
\gamma+e^{-} \rightarrow \gamma+e^{-} \quad \text { Implies } \\
e^{+}+e^{-} \rightarrow \gamma+\gamma \\
\gamma+\gamma \rightarrow e^{+}+e^{-}
\end{array}
\end{gathered}
$$

$$
\begin{gathered}
A+B \rightarrow C+D \rightarrow \quad \begin{array}{l}
\text { If this then } \\
\text { also... }
\end{array} \\
A \rightarrow \bar{B}+C+D \\
A+\bar{C} \rightarrow \bar{B}+D
\end{gathered}
$$

All that it tells us is whether there are any symmetry rules or conservation laws forbidding such a reaction (ie whether the dynamics are possible). Says nothing about the kinematics

$$
\begin{array}{ll}
A \rightarrow B+e^{-} & \begin{array}{l}
\text { What is the } \\
\text { energy of the } \\
\text { electron in this } \\
\text { decay? }
\end{array}
\end{array}
$$

## Neutrinos


$\mathrm{m}_{\mathrm{B}}$

$\mathrm{m}_{\mathrm{e}}$

Initially, in center
After decay of mass frame, everything is at rest, $\mathbf{p}=0$

$$
\begin{gathered}
E=m_{A} \\
p=0 \\
E=\sqrt{m_{B}^{2}+p_{B}^{2}}+\sqrt{m_{e}^{2}+p_{e}^{2}}
\end{gathered}
$$

Conservation of momentum: $\quad p_{B}=p_{e}=p$

## Neutrinos

$\mathrm{m}_{\mathrm{A}}$

## Initially, in center

 of mass frame, everything is at rest, $p=0$$$
\begin{gathered}
E=\sqrt{m_{B}^{2}+p^{2}}+\sqrt{m_{e}^{2}+p^{2}}=m_{A} \\
m_{B}^{2}+p^{2}+m_{e}^{2}+p^{2}+2 \sqrt{\left(m_{B}^{2}+p^{2}\right)\left(m_{e}^{2}+p^{2}\right)}=m_{A}^{2} \\
2 \sqrt{\left(m_{B}^{2}+p^{2}\right)\left(m_{e}^{2}+p^{2}\right)}=m_{A}^{2}-m_{B}^{2}-m_{e}^{2}-2 p^{2} \\
4\left(m_{B}^{2}+p^{2}\right)\left(m_{e}^{2}+p^{2}\right)=
\end{gathered}
$$

$$
m_{A}^{4}+m_{B}^{4}+m_{e}^{4}+4 p^{4}-2 m_{A}^{2} m_{B}^{2}-2 m_{A}^{2} m_{e}^{2}-4 m_{A}^{2} p^{2}+2 m_{B}^{2} m_{e}^{2}+4 m_{B}^{2} p^{2}+4 m_{e}^{2} p^{2}
$$

## Neutrinos


$\mathrm{m}_{\mathrm{B}}$

$\mathrm{m}_{\mathrm{e}}$

## Initially, in center

## is at rest, $p=0$

$$
\begin{gathered}
4 m_{B}^{2} m_{e}^{2}+4 m_{B}^{2} p^{2}+4 m_{e}^{2} p^{2}+4 p^{4}= \\
m_{A}^{4}+m_{B}^{4}+m_{e}^{4}+4 p^{4}-2 m_{A}^{2} m_{B}^{2}-2 m_{A}^{2} m_{e}^{2}-4 m_{A}^{2} p^{2}+2 m_{B}^{2} m_{e}^{2}+4 m_{B}^{2} p^{2}+4 m_{e}^{2} p^{2} \\
m_{A}^{4}+m_{B}^{4}+m_{e}^{4}-2 m_{A}^{2} m_{B}^{2}-2 m_{A}^{2} m_{e}^{2}-4 m_{A}^{2} p^{2}-2 m_{B}^{2} m_{e}^{2}=0 \\
p^{2}=\frac{1}{4}\left(m_{A}^{2}+\left(m_{B}^{4}+m_{e}^{4}-2 m_{B}^{2} m_{e}^{2}\right) / m_{A}^{2}-2 m_{B}^{2}-2 m_{e}^{2}\right)
\end{gathered}
$$

## Neutrinos

## Energy of electron is completely specified, but we observe a range of energies in neutron decay! Must be missing some object: neutrinos!

$$
\begin{gathered}
4 m_{B}^{2} m_{e}^{2}+4 m_{B}^{2} p^{2}+4 m_{e}^{2} p^{2}+4 p^{4}= \\
m_{A}^{4}+m_{B}^{4}+m_{e}^{4}+4 p^{4}-2 m_{A}^{2} m_{B}^{2}-2 m_{A}^{2} m_{e}^{2}-4 m_{A}^{2} p^{2}+2 m_{B}^{2} m_{e}^{2}+4 m_{B}^{2} p^{2}+4 m_{e}^{2} p^{2} \\
m_{A}^{4}+m_{B}^{4}+m_{e}^{4}-2 m_{A}^{2} m_{B}^{2}-2 m_{A}^{2} m_{e}^{2}-4 m_{A}^{2} p^{2}-2 m_{B}^{2} m_{e}^{2}=0 \\
p^{2}=\frac{1}{4}\left(m_{A}^{2}+\left(m_{B}^{4}+m_{e}^{4}-2 m_{B}^{2} m_{e}^{2}\right) / m_{A}^{2}-2 m_{B}^{2}-2 m_{e}^{2}\right)
\end{gathered}
$$

## Neutrinos

Energy of electron is completely specified, but we observe a range of energies in neutron decay! Must be missing some object: neutrinos!

## From Griffiths



Fig. 1.5 The beta decay spectrum of tritium $\left({ }_{1}^{3} \mathrm{H} \rightarrow{ }_{2}^{3} \mathrm{He}\right)$.
(Source: Lewis, C. M. (1970) Neutrinos, Wykeham, London,
p. 30.)

## "Little neutral one" to distinguish from the neutron. Physicists really cling to the idea of conservation of energy

## From Griffiths



Fig. 1.5 The beta decay spectrum of tritium $\left({ }_{1}^{3} \mathrm{H} \rightarrow{ }_{2}^{3} \mathrm{He}\right)$.
(Source: Lewis, C. M. (1970) Neutrinos, Wykeham, London,
p. 30.)

## Neutrinos, Cowan and Reines

Large tank of water mixed with cadmium chloride near a nuclear reactor to look for

$$
\begin{aligned}
& \bar{\nu}+p \rightarrow n+e^{+} \\
& e^{+}+e^{-} \rightarrow \gamma \gamma
\end{aligned}
$$

What about the neutron? It gets captured by cadmium, which then emits pairs of photons shortly thereafter

$$
n+\nu \rightarrow p+e^{-}
$$

$$
n+\bar{\nu} \nrightarrow p+e^{-}
$$

$$
\gamma+\gamma \rightarrow \mu^{+}+e^{-}
$$

$$
\mu^{-} \rightarrow e^{-}+\gamma
$$

Be careful about what is +1 and what is -1 !

Why not? Well, neutrinos and anti-neutrinos are not the same thing. But also... this violates a conversation law known as conservation of electron number. Electrons and electron neutrinos carry electron number $=+1$, and antielectrons and antielectron neutrinos $=-1$

Define in the same way a "muon number". Muons and muon neutrinos carry muon number $=+1$, and anti-muons and antimuon neutrinos $=-1$

And yet the same thing for
taus (which we haven't seen quite yet but we will soon)

|  | Before |  | After |  |
| :---: | :---: | :---: | :---: | :---: |
|  | e mu | e | mu |  |
| $\overline{\nu_{e}}+p \rightarrow n+e^{+}$ | -1 | 0 | -1 | 0 |
| $\pi^{+} \rightarrow \mu^{+}+\nu_{\mu}$ | 0 | 0 | 0 | $1-1=0$ |
| $\pi^{-} \rightarrow \mu^{-}+\overline{\nu_{\mu}}$ | 0 | 0 | 0 | $1-1=0$ |
| $\mu^{-} \rightarrow e^{-}+\overline{\nu_{e}}+\nu_{\mu}$ | 0 | 1 | $1-1=0$ | 1 |
| $n \rightarrow p+e^{-}+\overline{\nu_{e}}$ | 0 | 0 | $1-1=0$ | 0 |

Neutron decay

$$
n \rightarrow e^{-}+\overline{\nu_{e}}+p
$$

No proton decay

$$
p \nrightarrow e^{+}+\nu_{e}
$$

$$
p \rightarrow e^{+}+\nu_{e}+\gamma
$$

Proton decay is not observed! Why not? Propose conservation of baryon number $B=+1$ for neutrons and protons, -1 for antineutrons and anti-protons

Proton is lightest baryon, so it cannot decay (well, not in the Standard Model!)

No "conservation of meson" number

$$
K^{0} \rightarrow \pi^{+}+\pi^{-}
$$

## From Griffiths

Incident cosmicray
shower



Fig. 1.7 The first strange particle. Cosmic rays strike a lead plate, producing a $K^{0}$, which subsequently decays into a pair of charged pions. (Photo courtesy of Prof. Rochester, G. D. (© 1947). Nature, 160, 855. Copyright Macmillan Journals Limited.)

## What is this

 neutral kaon? Can produce them in accelerators, but they decay 13 orders of magnitude slower than expected! Strange...
## How to account for these strange particles?

What if kaons are produced by the strong force, but decay typically via the weak force?
Strangeness is conserved in strong interactions but not in weak interactions

$$
\begin{gathered}
\pi^{-}+p \rightarrow K^{+}+\Sigma^{-} \\
\pi^{-}+p \rightarrow K^{0}+\Sigma^{0} \\
\pi^{-}+p \rightarrow K^{0}+\Lambda
\end{gathered}
$$

Initially, have zero strangeness. Assign S=+1 to $\mathrm{K}^{+}$and $\mathrm{K}^{0}$ and $\mathrm{S}=-1$ to $\wedge$ and $\Sigma$, $S$ remains 0 . What does this say about how the above are produced? Note: strangeness not conserved in the decay.

## And how to start organizing all of this?



Murray Gell-Man proposed his "Eightfold way" (apparently a slight reference to the Noble Eightfold path of Buddhism)


Moving from baryons to mesons

$$
\begin{aligned}
& s=1 \\
& s=0 \\
& s=-1
\end{aligned}
$$

Similar periodic structure (remind you of anything)?

$$
\pi^{-} \pi^{K^{0}} \pi_{K^{0}}^{K^{0}} \pi^{K^{+}} \quad q=1
$$




We've skipped over some of the fun history

The particle ... zoo was becoming a big mess to keep track of

Wolfgang Pauli: "Had I foreseen that, I would have gone into botany"


Fermi to Lederman: "Young man, if I could remember the names of these particles, I would have been a botanist"

## We've skipped over some of the fun history

The particle ... zoo was becoming a big mess to keep track of (for some reason, physicists don't like botanists, apparently)



Three quarks for Muster Mark! Sure he has not got much of a bark And sure any he has it's all beside the mark.
—James Joyce, Finnegans Wake

## Quark and not kwork!

Wikipedia: Finnegans Wake is a novel by Irish writer James Joyce. It is significant for its experimental style and reputation as one of the most difficult works of fiction in the English language.

## FINNEGANS WAKE

## by

James Joyce

London
Faber and Faber Limited

## The quark model

## Hadrons



All hadrons (baryons and mesons) are made themselves of smaller pieces called quarks

|  | Charge (units of e) | Strangeness |
| :--- | :---: | :---: |
| u | $+2 / 3$ | 0 |
| d | $-1 / 3$ | 0 |
| s | $-1 / 3$ | 1 |
| ubar | $-2 / 3$ | 0 |
| dbar | $+1 / 3$ | 0 |
| sbar | $+1 / 3$ | -1 |

Mesons = one quark and one anti-quark Baryons = three quarks Anti-baryons = three anti-quarks

Example of the baryon decuplet

|  | Q | S | Baryon | B number |
| :--- | :---: | :---: | :---: | :---: |
| uuu | 2 | 0 | $\boldsymbol{\Delta}^{++}$ | 1 |
| uud | 1 | 0 | $\Delta^{+}$ | 1 |
| udd | 0 | 0 | $\boldsymbol{\Delta}^{0}$ | 1 |
| ddd | -1 | 0 | $\boldsymbol{\Delta}^{+}$ | 1 |
| uus | 1 | -1 | $\Sigma^{+}$ | 1 |
| uds | 0 | -1 | $\Sigma^{*}$ | 1 |
| dds | -1 | -1 | $\Sigma^{*}$ | 1 |
| uss | 0 | -2 | $\boldsymbol{\Xi}^{*} 0$ | 1 |
| dss | -1 | -2 | $\boldsymbol{\Xi}^{*}$ | 1 |
| sss | -1 | -3 | $\Omega^{-}$ | 1 |

Note that this fills in nicely and makes predictions, but we have yet to account for different energy levels and spin

See also the table of suggested $q \bar{q}$ quark－model assignments in the Quark Model section．
Indicates particles that appear in the preceding Meson Summary Table．We do not regard the other entries as being established．

|  | LIGHT UNFLAVORED$(S=C=B=0)$ |  |  | $\begin{gathered} \text { STRANGE } \\ (S= \pm 1, C=B=0) \end{gathered}$ |  | CHARMED，STRANGE$\begin{array}{r} (C=S= \pm 1) \\ \quad\left(J^{P}\right) \end{array}$ |  | ${ }^{C T}{ }_{1}{ }^{( }\left(J^{P C}\right)$ |  |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
|  | $I^{G}\left(J^{P C}\right)$ |  | $I^{G}\left(J^{P C}\right)$ |  | $1\left(J^{P}\right)$ |  |  |  | ） |
| －$\pi^{ \pm}$ | $1^{-}\left(0^{-}\right)$ | －$\phi$（1680） | $0^{-}\left(1^{--}\right)$ | －$K^{ \pm}$ | 1／2（0－） | －$D_{s}^{ \pm}$ | $0\left(0^{-}\right.$ | －J／$/ 4(15)$ | $0^{-}(1$ |
| －$\pi^{0}$ | $1^{-}\left(0^{-+}\right)$ | －$p_{3}(1690)$ | $1^{+}\left(3^{--}\right)$ | －$K^{0}$ | 1／2（0－） | －$D_{s}^{* \pm}$ | $0(?$ ？ | －$\chi_{\text {co }}(1 P)$ | $0^{+}\left(0^{++}\right)$ |
| －$\eta$ | $0^{+}\left(0^{-+}\right)$ | －$\rho(1700)$ | $1^{+}\left(1^{--}\right)$ | －$K_{S}^{0}$ | 1／2（0－） |  | $0\left(0^{+}\right)$ | －$\chi_{C 1}(1 P)$ | $0^{+}\left(1^{++}\right)$ |
| －$f_{0}(500)$ | $0^{+}\left(0^{++}\right)$ | $a_{2}(1700)$ | $1^{-}\left(2^{++}\right)$ | －$K_{L}^{0}$ | $1 / 2\left(0^{-}\right)$ |  | $0\left(1^{+}\right)$ | －$h_{c}(1 P)$ |  |
| －$\rho(770)$ | $1^{+}\left(1^{--}\right)$ | －$f_{0}(1710)$ | $0^{+}\left(0^{++}\right)$ | $K_{0}^{*}(800)$ | 1／2（ $0^{+}$） |  | $0\left(1^{+}\right)$ | －$\chi_{\text {c }}(1 P)$ | $0^{+}\left(2^{++}\right)$ |
| －$\omega$（782） | $0^{-}\left(1^{--}\right)$ | $\eta(1760)$ | $0^{+}\left(0^{-+}\right)$ | －$K^{*}$（892） | 1／2（ $1^{-}$） | －$D_{52}$ | $0(?$ ？$)$ | －$\eta_{c}(2 S)$ | $0^{+}\left(0^{-+}\right)$ |
| －$\eta^{\prime}(958)$ | $0^{+}\left(0^{-+}\right)$ | －$\pi(1800)$ | $1^{-}\left(0^{-+}\right)$ | －$K_{1}(1270)$ | 1／2（ $1^{+}$） | －$D_{s 1}^{*}$ | $0\left(1^{-}\right)$ | －$\psi(25)$ | $0^{-}\left(1^{--}\right)$ |
| －$f_{0}(980)$ | $0^{+}\left(0^{++}\right)$ | $\mathrm{f}_{2}(1810)$ | ${ }^{0}{ }^{+}\left(2^{++}\right)$ | －$K_{1}(1400)$ | 1／2（ $1^{+}$） |  | $0(? ?$ | －$\psi(3770)$ |  |
| －${ }^{\text {a }}$（980） | $1^{-\left(0^{++}\right)}$ | X $(1835)$ $X(1840)$ | $?^{?}$ ？？？？？？${ }^{+}$） | － $\mathrm{K}^{*}(1410)$ | $1 / 2\left(1^{-}\right)$ $1 / 20^{+}$ |  | $0(?)$ | $X(3823)$ $-\times(3872)$ |  |
| －$\phi(1020)$ <br> －$h_{1}(1170)$ | $\begin{aligned} & 0^{-}\left(1^{--}\right) \\ & 0^{-}\left(1^{+-}\right) \end{aligned}$ | $X(1840)$ －${ }_{3}(1850)$ | $? ?(? \cdot)$ $0^{-}\left(3^{-}\right)$ | －${ }^{-K_{0}^{*}(1430)}$ | $\begin{aligned} & 1 / 2\left(0^{+}\right) \\ & 1 / 2\left(2^{+}\right) \end{aligned}$ |  |  | － $\begin{aligned} & \text {－} \times(38872) \\ & \text {（3900）} \\ & \\ & \text {（ }\end{aligned}$ | $\begin{aligned} & 0^{+}\left(1^{++}+\right. \\ & ?\left(1^{+}\right) \end{aligned}$ |
| －$b_{1}(1235)$ | $1^{+}\left(1^{+-}\right)$ | $\eta_{2}(1870)$ | $0^{+}\left(2^{-+}\right)$ | $\begin{aligned} & \bullet K_{2}^{2}(1430) \\ & K(1460) \end{aligned}$ | $\begin{aligned} & 1 / 2\left(2^{\top}\right) \\ & 1 / 2\left(0^{-}\right) \end{aligned}$ |  |  | X（3900）${ }^{0}$ | ？（？？） |
| －$a_{1}(1260)$ | $1^{-}\left(1^{++}\right)$ | －$\pi_{2}(1880)$ | $1^{-}\left(2^{-+}\right)$ | $K_{2}(1580)$ | 1／2（2－） | －$B^{ \pm}$ | 1／2（0） | －$\chi_{c o}(2 P)$ | $0^{+}\left(0^{++}\right)$ |
| －$f_{2}(1270)$ | $0^{+}\left(2^{++}\right)$ | $\rho(1900)$ | $1^{+}\left(1^{--}\right)$ | K（1630） | 1／2（？？） | －$B^{0}$ | 1／2（0－） | －$\chi_{\text {c } 2}(2 P)$ | ${ }^{0+}(2++)$ |
| －$f_{1}(1285)$ | $0^{+}\left(1^{++}\right)$ | $f_{2}(1910)$ | $0^{+}\left(2^{++}\right)$ | $K_{1}(1650)$ | 1／2（ $1^{+}$） | －$B^{ \pm} / B^{\text {a }}$ | IXTURE | ${ }^{X(3940)}$ | ？？（？？？ |
| －$\eta(1295)$ | $0^{+}\left(0^{-+}\right)$ | －$f_{2}(1950)$ | $0^{+}\left(2^{++}\right)$ | －$K^{*}(1680)$ | $1 / 2\left(1^{-}\right)$ | －$B^{ \pm}$ | －baryon | $x(4020)^{ \pm}$ | ？（？？） |
| －$\pi(1300)$ | $1^{-}\left(0^{-+}\right)$ | $\rho_{3}(1990)$ | $1^{+}\left(3^{--}\right)$ | －$K_{2}(1770)$ | 1／2（2）${ }^{-}$ |  |  | －$\psi(4040)$ | $0^{-}\left(1^{--}\right)$ |
| －$a_{2}(1320)$ | $1^{-}\left(2^{++}\right)$ | －$f_{2}(2010)$ | $0^{+}\left(2^{++}\right)$ | －$K_{3}^{*}(1780)$ | 1／2（3） |  | CKM Ma－ | $\left.{ }^{X(4050}\right)^{ \pm}$ | ？？（？${ }^{\text {a }}$（ ${ }^{\text {a }}$ |
| －$f_{0}(1370)$ | $0^{+}\left(0^{++}\right)$ | $f_{0}(2020)$ | $0^{+}\left(0^{++}\right)$ | －$K_{2}(1820)$ | 1／2（2－） | －$B^{*}$ | 1／2（1－） | $X(4140)$ | ${ }^{0+}\left(?^{++}\right)$ |
| $h_{1}(1380)$ | $?^{-\left(1^{+-}\right)}$ | － $\mathrm{a}_{4}(2040)$ | $1^{-}(4++)$ | K（1830） | 1／2（ $0^{-}$） |  | ？（？？） | －$\psi(4160)$ | $0^{-}\left(1{ }^{--}\right.$ |
| －$\pi_{1}(1400)$ | $1^{-}\left(1^{-+}\right)$ | －$f_{4}(2050)$ | $0^{+}\left(4^{++}\right)$ | $K_{0}^{*}(1950)$ | $1 / 2\left(0^{+}\right)$ | －$B_{1}(5$ | 1／2（1＋） | $X(4160)$ $X(4250)^{ \pm}$ | $\left.?_{?}^{? ?(? ? ?}\right)$ |
| －$\eta(1405)$ | $0^{+}\left(0^{-+}\right)$ | $\pi_{2}(2100)$ | $\mathrm{l}^{-\left(2^{-+}\right)}$ | $K_{2}^{*}(1980)$ | 1／2（2＋） | －$B_{2}^{*}(5$ | 1／2（2＋） | $X(4250)^{ \pm}$ <br> －$X(4260)$ | $\begin{aligned} & ? ? ?(?) \\ & ? ?(1--) \end{aligned}$ |
| －$f_{1}(1420)$ | $0^{+}\left(1^{++}\right)$ | $\mathrm{f}_{0}(2100)$ | $0^{+}\left(0^{++}\right)$ | －$K_{4}^{*}(2045)$ | 1／2（4＋） |  |  | －$X(4260)$ <br> X（4350） | $\begin{aligned} & ? ?(1--) \\ & ?(?++) \end{aligned}$ |
| －$\omega$（1420） | $0^{-}\left(1^{--}\right)$ $0^{+}\left(2^{++}\right)$ | $f_{2}(2150)$ $\rho(2150)$ | $0^{+}(2++)$ $1^{+}\left(1^{--}\right)$ | －${ }^{\text {K }}$ 2（2250） | 1／2（2－） |  | $\begin{aligned} & \text { TRANGE } \\ & =\mp \mp 1) \end{aligned}$ | $\begin{array}{r} X(4350) \\ -X(4360) \end{array}$ | $\begin{aligned} & 0^{+}+?^{?+}+ \\ & ?^{?}\left(1^{--}\right) \end{aligned}$ |
| － $0_{0}(1450)$ | $1^{-}\left(0^{++}\right)$ | －$\phi(2170)$ | $0^{-}\left(1^{--}\right)$ | $K_{3}(2320)$ $K^{*}(2380)$ | 1／2（3＋） | －$B_{s}^{0}$ | $0\left(0^{-}\right)$ | －$\psi(4415)$ | $0^{-}\left(1^{--}\right)$ |
| －$\rho(1450)$ | $1^{+}\left(1^{--}\right)$ | $\mathrm{f}_{0}(2200)$ | $0^{+}\left(0^{++}\right)$ | $K_{5}^{*}(2380)$ $K_{4}(2500)$ |  |  | $0\left(1^{-}\right)$ | $X(4430)^{ \pm}$ | ？${ }^{(1+}$ ） |
| －$\eta(1475)$ | $0^{+}\left(0^{-+}\right)$ | $f_{J}(2220)$ | $0^{+}\left(2^{++}\right.$ | $\begin{aligned} & K_{4}(2500) \\ & K(3100) \end{aligned}$ | $\begin{aligned} & 1,2\left(4^{-}\right) \\ & ? ?(? ? ?) \end{aligned}$ |  | $0\left(1^{+}\right)$ | －X（4660） | ） |
| －$f_{0}(1500)$ | $0^{+}\left(0^{++}\right)$ |  | or $4++$ $0^{+}$ |  |  | －$B_{52}^{*}$ | O（2＋） | $b \bar{b}$ |  |
| ${ }^{f_{1}(1510)}$ | $0^{+}\left(1^{++}\right)$ | $\eta$（2225） | $0^{0+}\left(0^{-+}\right)$ | $\begin{aligned} & \text { CHAR } \\ & (C=1 \end{aligned}$ |  |  | ？（？${ }^{\text {？}}$ ） |  |  |
| －$f_{2}^{\prime}(1525)$ | $0^{+}\left(2^{++}\right)$ | $\rho_{3}(2250)$ | $1^{+}\left(3^{--}\right)$ |  |  |  |  | $\eta_{b}(1 S)$ $\text { - } r(1 S)$ | $\begin{aligned} & 0^{+}\left(0^{-+}\right) \\ & 0^{-}\left(1^{---}\right) \end{aligned}$ |
| $\mathrm{f}_{2}(1565)$ | $0^{+}\left(2^{++}\right)$ | －$f_{2}(2300)$ | $0^{+}\left(2^{++}\right)$ |  | $1 / 2\left(0^{-}\right)$ |  | HARMED $\pm 1)$ | －$r(1 S)$ <br> －$\chi_{\infty 0}(1 P)$ | $\begin{aligned} & 0^{-}\left(1^{--}\right) \\ & 0^{+}\left(0^{++}\right. \end{aligned}$ |
| $\rho(1570)$ | $1^{+}\left(1^{--}\right)$ | $f_{4}(2300)$ | $0^{+}\left(4^{++}\right)$ | －$D^{0}$ | $1 / 2\left(0^{-}\right)$ |  |  | －$\chi_{\text {bo }}(1 P)$ <br> －$\chi_{b 1}(1 P)$ | $\begin{array}{r} +\left(0^{++}\right) \\ +\left(1^{++}\right) \end{array}$ |
| $h_{1}(1595)$ | $0^{-}\left(1^{+-}\right)$ | $f_{0}(2330)$ | $0^{+}\left(0^{++}\right)$ | －$D^{*}(2007)^{0}$ | 1／2（1－） | － | $0\left(0^{-}\right)$ | －$\chi_{b 1}(1 P)$ <br> －$h_{b}(1 P)$ | $\begin{aligned} & 0_{?}^{+}+\left(1^{++}+-\right) \end{aligned}$ |
| －$\pi_{1}(1600)$ | $1^{-}\left(1^{-+}\right)$ | －$f_{2}(2340)$ | $0^{+}\left(2^{++}\right)$ | －D ${ }^{*}(2010)^{ \pm}$ | $1 / 2\left(1^{-}\right)$ |  |  | －$h_{b}(1 P)$ <br> －$\chi_{n 2}(1 P)$ | $\begin{aligned} & ? ?(1+-) \\ & 0^{+}(0++ \end{aligned}$ |
| $a_{1}(1640)$ | $1^{-}\left(1^{++}\right)$ | $\rho_{5}(2350)$ | $1^{+}\left(5^{--}\right)$ | $\text { - } D_{0}^{+}(2400)^{0}$ | $1 / 2\left(0^{+}\right)$ |  |  | －$\chi_{b 2}(1 P)$ | $\begin{aligned} & 0^{+}\left(2^{++}\right) \\ & 0^{+}\left(0^{-+}\right. \end{aligned}$ |
| $\mathrm{f}_{2}(1640)$ | $0^{+}\left(2^{++}\right)$ | $a_{6}(2450)$ | $1^{-}\left(6^{++}\right)$ | $D_{0}^{*}(2400)^{ \pm}$ | 1／2（ $0^{+}$） |  |  | $\eta_{b}(2 S)$ | $0^{+}\left(0^{-+}\right)$ |
| －$\eta_{2}(1645)$ | $0^{+}\left(2^{-+}\right)$ | $\mathrm{f}_{6}(2510)$ | $0^{+}\left(6^{++}\right)$ | －$D_{1}(2420)^{0}$ | 1／2（ $1^{+}$） |  |  | $\text { - } r(2 S)$ <br> －$r(1 D)$ | $\begin{aligned} & 0^{-}\left(1^{--}\right) \\ & 0^{-}\left(2^{--}\right) \end{aligned}$ |
| －$\omega$（1650） | $0^{-}(1$ | OTHER | LIGHT | $D_{1}(2420)^{ \pm}$ | 1／2（？？） |  |  | －$r(1 D)$ <br> －$\chi_{\text {oo }}(2 P)$ | $\begin{aligned} & 0^{-}\left(2^{--}\right) \\ & 0^{+}\left(0^{+}\right) \end{aligned}$ |
| －${ }^{\text {－}} \omega_{3}(16670)$ | $0^{-}{ }^{-}\left(3^{-}\right.$ | Further St |  | $D_{1}(2430)^{0}$ | 1／2（1＋） |  |  | －$\chi_{b 1}(2 P)$ | ${ }^{+}+(1++)$ |
|  |  |  |  | －$D_{2}^{*}(2460)^{0}$ | 1／2（2＋） |  |  | $h_{b}(2 P)$ | $? ?\left(1^{+-}\right)$ |
|  |  |  |  | －$D_{2}^{*}(2460)^{ \pm}$ | 1／2（2＋） |  |  | $\cdot \chi_{b 2}(2 P)$ | $0^{+}\left(2^{++}\right)$ |
|  |  |  |  | $D(2550)^{0}$ | 1／2（0） |  |  | $\text { - } r(3 S)$ |  |
|  |  |  |  | $D(2600)$ | 1／2（？？） |  |  | $\text { - } \chi \text { b }(3 P)$ | $\text { ?? } ?\left(?^{?+}\right)$ |
|  |  |  |  | $D^{*}(2640){ }^{ \pm}$ | 1／2（？？${ }^{\text {a }}$ ） |  |  | －$r(4 S)$ | $0^{-}\left(1^{--}\right)$ |
|  |  |  |  | $D(2750)$ | 1／2（？？） |  |  | $x(10610)^{ \pm}$ | $1^{+}\left(1^{+}\right)$ |
|  |  |  |  |  |  |  |  | $X(10610)^{0} 1$ | $1^{+}\left(1^{+}\right)$ |
|  |  |  |  |  |  |  |  | $X(10650)^{ \pm}$？ | $?^{+}\left(1^{+}\right)$ |
|  |  |  |  |  |  |  |  | －$r(10860)$ | $0^{-}\left(1^{--}\right)$ |
|  |  |  |  |  |  |  |  | －$r(11020)$ | $0^{-}\left(1^{-}\right.$ |

Baryon Summary Table
This short table gives the name，the quantum numbers（where known），and the status of baryons in the Review．Only the baryons with 3－or 4－star status are included in the Baryon Summary Table．Due to insufficient data or uncertain interpretation，the other entries in the table
are not established baryons．The names with masses are of baryons that decay strongly．The spin－parity $J^{P}$（when known）is given with each particle．For the strongly decaying particles，the $J^{P}$ values are considered to be part of the names．

| $p$ | $1 / 2^{+}$ | ＊＊＊＊ | $\Delta(1232)$ | $3 / 2^{+}$ | ＊＊＊＊ | $\Sigma^{+}$ | 1／2＋ | ＊＊＊＊ | 三 | $1 / 2^{+}$ | ＊＊＊＊ | $\Lambda_{c}^{+}$ | $1 / 2^{+}$ | ＊＊＊＊ |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
| $n$ | $1 / 2^{+}$ | ＊＊＊＊ | $\Delta(1600)$ | $3 / 2^{+}$ | ＊＊＊ | $\Sigma^{0}$ | 1／2 ${ }^{+}$ | ＊＊＊ | 三－ | $1 / 2^{+}$ | ＊＊＊＊ | $\Lambda_{c}(2595)^{+}$ | 1／2－ | ＊＊＊ |
| $N(1440)$ | $1 / 2^{+}$ | ＊＊＊＊ | $\Delta(1620)$ | $1 / 2^{-}$ | ＊＊＊＊ | $\Sigma$ | 1／2 ${ }^{+}$ | ＊＊＊＊ | 三（1530） | 3／2＋ | ＊＊＊＊ | $\Lambda_{c}(2625)^{+}$ | $3 / 2^{-}$ | ＊＊＊ |
| $N(1520)$ | 3／2 ${ }^{-}$ | ＊＊＊＊ | $\Delta(1700)$ | $3 / 2^{-}$ | ＊＊＊＊ | $\Sigma(1385)$ | $3 / 2^{+}$ | ＊＊＊＊ | 三（1620） |  | ＊ | $\Lambda_{c}(2765)^{+}$ |  | ＊ |
| $N(1535)$ | $1 / 2^{-}$ | ＊＊＊＊ | $\Delta(1750)$ | $1 / 2^{+}$ | ＊ | $\Sigma(1480)$ |  | ＊ | 三（1690） |  | ＊＊＊ | $\Lambda_{c}(2880)^{+}$ | $5 / 2^{+}$ | ＊＊＊ |
| $N(1650)$ | $1 / 2^{-}$ | ＊＊＊＊ | $\Delta(1900)$ | $1 / 2^{-}$ | ＊＊ | $\Sigma(1560)$ |  | ＊＊ | 三（1820） | $3 / 2^{-}$ | ＊＊＊ | $\Lambda_{C}(2940)^{+}$ |  | ＊＊＊ |
| $N(1675)$ | 5／2－ | ＊＊＊ | $\Delta(1905)$ | 5／2＋ | ＊＊＊＊ | $\Sigma(1580)$ | $3 / 2^{-}$ | ＊ | 三（1950） |  | ＊＊＊ | $\Sigma_{c}(2455)$ | $1 / 2^{+}$ | ＊＊＊＊ |
| $N(1680)$ | $5 / 2^{+}$ | ＊＊＊＊ | $\Delta(1910)$ | $1 / 2^{+}$ | ＊＊＊＊ | $\Sigma(1620)$ | $1 / 2^{-}$ | ＊ | 三（2030） | $\geq \frac{5}{2}$ ？ | ＊＊＊ | $\Sigma_{c}(2520)$ | $3 / 2^{+}$ | ＊＊＊ |
| $N(1685)$ |  | ＊ | $\Delta$（1920） | $3 / 2^{+}$ | ＊＊＊ | $\Sigma(1660)$ | $1 / 2^{+}$ | ＊＊＊ | 三（2120） |  | ＊ | $\Sigma_{c}(2800)$ |  | ＊＊＊ |
| $N(1700)$ | $3 / 2^{-}$ | ＊＊＊ | $\Delta(1930)$ | 5／2－ | ＊＊＊ | $\Sigma(1670)$ | $3 / 2^{-}$ | ＊＊＊＊ | 三（2250） |  | ＊＊ | $\bar{E}_{c}^{+}$ | $1 / 2^{+}$ | ＊＊＊ |
| $N(1710)$ | $1 / 2^{+}$ | ＊＊＊ | $\Delta(1940)$ | $3 / 2^{-}$ | ＊＊ | $\Sigma(1690)$ |  | ＊＊ | 三（2370） |  | ＊＊ | ${ }^{\circ} \mathrm{c}$ | 1／2 ${ }^{+}$ | ＊＊＊ |
| $N(1720)$ | $3 / 2^{+}$ | ＊＊＊＊ | $\Delta(1950)$ | 7／2＋ | ＊＊＊＊ | $\Sigma(1730)$ | $3 / 2^{+}$ | ＊ | 三（2500） |  | ＊ | $\Xi_{c}^{\prime \prime}$ | $1 / 2^{+}$ | ＊＊＊ |
| $N(1860)$ | 5／2＋ | ＊＊ | $\Delta(2000)$ | 5／2＋ | ＊＊ | $\Sigma(1750)$ | 1／2 ${ }^{-}$ | ＊＊＊ |  |  |  | $\bar{E}^{\prime \prime}$ | 1／2 ${ }^{+}$ | ＊＊＊ |
| $N(1875)$ | $3 / 2^{-}$ | ＊＊＊ | $\Delta(2150)$ | $1 / 2^{-}$ | ＊ | $\Sigma(1770)$ | $1 / 2^{+}$ | ＊ | $\Omega^{-}$ | $3 / 2^{+}$ | ＊＊＊ | －${ }^{c}$（2645） | $3 / 2^{+}$ | ＊＊＊ |
| $N(1880)$ | 1／2 ${ }^{+}$ | ＊＊ | $\Delta(2200)$ | 7／2－ | ＊ | $\Sigma(1775)$ | 5／2－ | ＊＊＊＊ | $\Omega(2250)^{-}$ |  | ＊＊ | $\bar{\Xi}_{c}(2790)$ | $1 / 2^{-}$ | ＊＊＊ |
| $N(1895)$ | $1 / 2^{-}$ | ＊＊ | $\Delta(2300)$ | 9／2 ${ }^{+}$ | ＊＊ | $\Sigma(1840)$ | $3 / 2^{+}$ | ＊ | $\Omega(2380)^{-}$ |  | ＊＊ | $\bar{\Xi}_{c}(2815)$ | $3 / 2^{-}$ | ＊＊＊ |
| $N(1900)$ | $3 / 2^{+}$ | ＊＊＊ | $\Delta(2350)$ | 5／2－ | ＊ | $\Sigma(1880)$ | $1 / 2^{+}$ | ＊＊ | $\Omega(2470)^{-}$ |  | ＊＊ | $\bar{\Xi}_{c}(2930)$ |  | ＊ |
| $N(1990)$ | 7／2＋ | ＊＊ | $\Delta(2390)$ | 7／2＋ | ＊ | $\Sigma(1900)$ | $1 / 2^{-}$ | ＊ |  |  |  | $\bar{\Xi}_{c}(2980)$ |  | ＊＊＊ |
| $N(2000)$ | 5／2＋ | ＊＊ | $\Delta(2400)$ | $9 / 2^{-}$ | ＊＊ | $\Sigma(1915)$ | 5／2＋ | ＊＊＊＊ |  |  |  | $\bar{\Xi}_{c}(3055)$ |  | ＊＊ |
| $N(2040)$ | $3 / 2^{+}$ | ＊ | $\Delta(2420)$ | $11 / 2^{+}$ | ＊＊＊＊ | $\Sigma(1940)$ | $3 / 2^{+}$ | ＊ |  |  |  | $\bar{E}_{c}(3080)$ |  | ＊＊＊ |
| $N(2060)$ | 5／2－ | ＊＊ | $\Delta(2750)$ | 13／2－ | ＊＊ | $\Sigma(1940)$ | $3 / 2^{-}$ | ＊＊＊ |  |  |  | $\bar{E}_{c}(3123)$ |  | ＊ |
| $N(2100)$ | $1 / 2^{+}$ | ＊ | $\Delta(2950)$ | 15／2＋ | ＊＊ | $\Sigma(2000)$ | $1 / 2^{-}$ | ＊ |  |  |  |  | $1 / 2^{+}$ | ＊＊＊ |
| $N(2120)$ | $3 / 2^{-}$ | ＊＊ |  |  |  | $\Sigma(2030)$ | 7／2＋ | ＊＊＊＊ |  |  |  | $\Omega_{c}(2770)^{0}$ | $3 / 2^{+}$ | ＊＊＊ |
| $N(2190)$ | 7／2－ | ＊＊＊＊ | 1 | $1 / 2^{+}$ | ＊＊＊＊ | $\Sigma(2070)$ | 5／2＋ | ＊ |  |  |  | $\Omega_{c}(2770)$ |  |  |
| $N(2220)$ | 9／2 ${ }^{+}$ | ＊＊＊＊ | $\wedge$（1405） | 1／2－ | ＊＊＊＊ | $\Sigma(2080)$ | $3 / 2^{+}$ | ＊＊ |  |  |  | $\Xi_{c c}^{+}$ |  | ＊ |
| $N(2250)$ | 9／2－ | ＊＊＊＊ | $\wedge$＾1520） | $3 / 2^{-}$ | ＊＊＊＊ | $\Sigma(2100)$ | $7 / 2^{-}$ | ＊ |  |  |  |  |  |  |
| $N(2300)$ | $1 / 2^{+}$ | ＊＊ | $\wedge$＾1600） | $1 / 2^{+}$ | ＊＊＊ | $\Sigma(2250)$ |  | ＊＊＊ |  |  |  |  | 1／2 ${ }^{+}$ | ＊＊＊ |
| $N(2570)$ | 5／2 ${ }^{-}$ | ＊＊ | $\wedge$＾1670） | $1 / 2^{-}$ | ＊＊＊＊ | $\Sigma(2455)$ |  | ＊＊ |  |  |  | $\Lambda_{b}(5912)^{0}$ | 1／2 ${ }^{-}$ | ＊＊＊ |
| $N(2600)$ | 11／2 ${ }^{-}$ | ＊＊＊ | ＾（1690） | 3／2－ | ＊＊＊＊ | $\Sigma(2620)$ |  | ＊＊ |  |  |  | $\Lambda_{b}(5920)^{0}$ | 3／2－ | ＊＊＊ |
| $N(2700)$ | $13 / 2^{+}$ |  | $\wedge$（1710） | $1 / 2^{+}$ | ＊ | $\Sigma(3000)$ |  | ＊ |  |  |  |  | 1／2 ${ }^{+}$ | ＊＊＊ |
|  |  |  | $\wedge$＾（1800） | $1 / 2^{-}$ | ＊＊＊ | $\Sigma(3170)$ |  | ＊ |  |  |  |  |  | ＊＊＊ |
|  |  |  | $\wedge$＾1810） | $1 / 2^{+}$ | ＊＊＊ |  |  |  |  |  |  | $\overline{E S}_{b}^{0}, \bar{E}_{b}^{-}$ | $1 / 2^{+}$ | ＊＊＊ |
|  |  |  | $\Lambda(1820)$ $\Lambda(1830)$ | 5／2＋${ }^{+}$ | ${ }^{* * * *}$ |  |  |  |  |  |  | $\Xi_{b}(5945)^{0}$ | $3 / 2^{+}$ | ＊＊＊ |
|  |  |  | $\begin{aligned} & \Lambda(1830) \\ & \Lambda(1890) \end{aligned}$ |  | ${ }^{* * * * *}$ |  |  |  |  |  |  |  |  |  |
|  |  |  | $\wedge$（2000） |  | ＊ |  |  |  |  |  |  |  |  |  |
|  |  |  | $\wedge$（2020） | 7／2 ${ }^{+}$ | ＊ |  |  |  |  |  |  |  |  |  |
|  |  |  | $\wedge$（2050） | $3 / 2^{-}$ | ＊ |  |  |  |  |  |  |  |  |  |
|  |  |  | $\wedge$（2100） | 7／2－ | ＊＊＊＊ |  |  |  |  |  |  |  |  |  |
|  |  |  | $\wedge$（2110） | 5／2＋ | ＊＊＊ |  |  |  |  |  |  |  |  |  |
|  |  |  | $\wedge$（2325） | 3／2 ${ }^{-}$ | ＊ |  |  |  |  |  |  |  |  |  |
|  |  |  | $\Lambda(2350)$ | 9／2＋ | ＊＊＊ |  |  |  |  |  |  |  |  |  |
|  |  |  | 1 （2585） |  |  |  |  |  |  |  |  |  |  |  |

Existence is certain，and properties are at least fairly well explored．
＊＊Existence ranges from very likely to certain，but further confirmation is desirable and／or

Evidence of existence is only fair．
＊Evidence of existence is poor．

Quarks are confined to hadrons, and cannot be observed as bare objects. In other words, they are social beasts, and don't like to exist on their own. We'll discuss them more later on

## What about Pauli exclusion principle?

## sss hadron ( $\Omega^{-}$)

 should violate Pauli exclusion principle (multiple times!) There needs to be some fundamental difference between the quarks for this to be allowed: QCD color charge (beware comparisons to real "color")Fig. 1.11 (a) tn Rutherford scattering, the number of particles deflected through large angles indicates that the atom has internal structure (a nucleus). (b) in deep inelastic scattering, the number of particles deflected through large angles indicates that the proton has internal structure (quarks). The dashed lines show what you would expect
if the positive charge were uniformly distributed over the volume of (a) the atom, (b) the proton. (Source: Halzen, F. and Martin, A. D. (1984) Quarks and Leptons, John Wiley \& Sons, New York, p. 17. Copyright © john Wiley \& Sons, Inc. Reprinted by permission.)

## Colorless objects only: Explains why we don't have qq (or q) final states

## November revolution

November 1974, $\psi$ meson discovered at SLAC, J meson at Brookhaven. Hence the name $\mathrm{J} / \psi$, a new, electrically neutral particle with a long lifetime. Began the "November revolution"

iViva la Revolución!


## https://www.bnl.gov/ bnlweb/history/nobel/ nobel 76.asp

Sam Ting and his team, showing a plot of mass of $e^{+} e^{-}$pairs. As we will see shortly, the mass of objects is an invariant even after decay, so we see a bump at 3.1 GeV = JPsi mass

## And these days



# We know that the JPsi is a charmanticharm bound state with $\mathrm{spin}=1$ 

This is a tiny fraction of the data available to ATLAS and CMS, and arXiv: 1612.02950 a large fraction of the JPsi that we produced are not even recorded

## JPsi elsewhere



## Add in the third generation

Table I. From Perl (1975). A table of 2-charged-particle events collected at 4.8 GeV in the Mark I detector. The table, containing $24 e \mu$ events with zero total charge and no photons, was the strongest evidence at that time for the $\tau$. The caption read:
"Distribution of $513,4.8 \mathrm{GeV}$, 2-prong, events which meet the criteria: $p_{e}>0.65 \mathrm{GeV} / \mathrm{c}, p_{\mu}>0.65 \mathrm{GeV} / \mathrm{c}, \theta_{\text {copl }}>20^{\circ}$."

|  | Total Charge $=0$ |  |  | Total Charge $= \pm 2$ |  |  |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: |
| Number photons $=$ | 0 | 1 | $>1$ | 0 | 1 | $>1$ |
| $c c$ | 40 | 111 | 55 | 0 | 1 | 0 |
| $e \mu$ | 24 | 8 | 8 | 0 | 0 | 3 |
| $\mu \mu$ | 16 | 15 | 6 | 0 | 0 | 0 |
| $e h$ | 18 | 23 | 32 | 2 | 3 | 3 |
| $\mu h$ | 15 | 16 | 31 | 4 | 0 | 5 |
| $h h$ | 13 | 11 | 30 | 10 | 4 | 6 |
| Sum | 126 | 184 | 162 | 16 | 8 | 17 |



## Martin Perl (1975) discovery of tau lepton (tau neutrino not until 2000!)

Lederman and E288 collaborators - discovery of the bottom quark at nearby Fermilab:

$$
\Upsilon=b \bar{b}
$$



## Lots of upsilons everywhere, too

arxiv: 1804.09214 We'll talk about the 3 signal peaks in a few chapters!


## LHCb-PAPER-2018-035

LHCb is an asymmetric detector (any guesses why?). Difference between the two is which beams moves toward the detector at the interaction region



Announced jointly 1994-1995 by DZero and CDF (at Fermilab) in proton-antiproton collisions, but top quarks decay so quickly that they do not
form stable/metastable bound states. Mass of top quark $\sim 173 \mathrm{GeV}$ !


Fun aside: UA1 at CERN made a "discovery" of the top quark in 1984 with a mass of 40 +/- 10 GeV

## Fermions Bosons



This is really the full Standard Model (modulo the Higgs boson, h), though it of course ... hides some details

Are there additional generations of quarks? We've been looking for them

Three generations of quarks, with each generation getting more massive. Each quark carries electric charge (+2/3,-1/3), and also QCD color (rgb). Quarks are confined, and do not exist alone in nature, but rather only in hadrons: baryons and mesons. Quarks also contain "flavor" (strangeness, topness, etc) that is conserved in QCD, but not in weak interactions. Anti-quarks not shown


Are there additional leptons? We've been looking for them, though, though extra neutrinos have to be very massive or rarely interact. Masses of observed neutrinos unknown, but are very small

Three generations of leptons, with each generation (at least of charged leptons) getting more massive. Charged leptons carry electric charge, neutrinos do not. All leptons carry electron, muon and tau number, which is conserved. Anti-particles not shown

## Study

number of light
neutrinos with Z production at $\mathrm{e}^{+} \mathrm{e}^{-}$machines

## Photon never changes matter flavor



Force carriers are bosons with integer spin. The photon is the force carrier from
E\&M, and has spin-1
(it's a vector boson) and zero mass.
Interacts only with electrically charged particles

## The bosons

## Gluon never changes matter flavor



The gluon is the force carrier of QCD, and has spin 1 (it's a vector boson) and zero mass. It has no electric charge, but it carries QCD color charge (there are 8 types of gluons). Hence, it couples not only to quarks, but also to itself and other gluons (but not leptons)



## W bosons

always change matter flavor,
$Z$ bosons do not

The $\mathrm{W}^{ \pm}$and $\mathrm{Z}^{0}$ bosons are the weak force carriers, with spin-1 (vector bosons) and both having large, non-zero mass (W~80 $\mathrm{GeV}, \mathrm{Z} \sim 91 \mathrm{GeV}$ ). The weak force is called exactly that due to their large mass. Weak force carriers are special - they are the only way to change one quark generation into another (see JPsi lifetime!), and have other special properties that we'll get to. Responsible for nuclear $\beta$ decay and fusion

The Higgs boson (h) was hypothesized in 1960s, but was not discovered until July 2012. Long timeline! We'll see later in the course how the Higgs mechanism explains why the weak force carriers are so weak/ massive, and why the fermions have mass

$$
\mathrm{mH} \sim 125 \mathrm{GeV}
$$

arXiv: 1902.05892

## 13 TeV (moving likely to 14 TeV ) protonproton collisions produced by the LHC



CERN<br>$\rightarrow$ ATLAS<br>Point 1

ALICE
$\Rightarrow$ Point 2


## ATLAS (I can talk about my experiment for.. awhile)



# Trigger system: 40 MHz 

## ATLAS

# reduced to $\sim 1 \mathrm{kHz}$ 

# reduced to $\sim 1 \mathrm{kHz}$ 














Producing collisions



$\qquad$







## g-2

mu2e
neutrino experiments (nuclear, solar, accelerator) nuclear physics
astroparticle physics
dark matter searches
dark energy surveys

Not to mention all the accelerator physics that goes along with much of the above

Want to see the effect of a particle, usually in the form of radiation or energy transfer that we can observe. Goal is to infer as much about the original particle/object as possible.

Problem is of course that we cannot just put things inside of a microscope for careful, lengthy study: typically traveling close to speed of light, often very short-lived, occasionally produced with many other nearby objects, sometimes have no electric charge and can look like other, similar particles

## What can we take advantage of?

Charged particles bend in a magnetic field (can measure electric charge and momentum)

Careful timing measurements can give us particle velocity and other information (useful not only for the original particle, but also secondary particles it produce)

Calorimeter measurements give particle energy (different types of particles deposit energy electromagnetically vs via strong force, and at different rates)

Neutrinos escape $\sim$ undetected, leading to imbalance of momentum in detector

Muons are highly penetrating and minimally ionizing
Other tricks like Cherenkov radiation, measurements of $\mathrm{dE} / \mathrm{dx}$, looking for transition radiation at material boundaries, ...



玉
흥
응


100 million readout channels (upgraded between the two data-taking years on the plot for an extra layer) to measure precise location of charged particle trajectory
https://atlas.web.cern.ch/Atlas/GROUPS/ PHYSICS/PLOTS/IDTR-2015-007/

## Semiconductors



## The $p-n$ Junction as a Tracking Detector

- thin $(\sim \mu \mathrm{m})$, highly doped $p^{+}\left(\sim 10^{19} \mathrm{~cm}^{-3}\right)$ layer on lightly doped $n\left(\sim 10^{12} \mathrm{~cm}^{-3}\right)$ substrate
- high mobility of charge carriers in Si allows fast charge collection ( $\sim 5 \mathrm{~ns}$ for electron)
- high Si density \& low electron-hole creation potential ( 3.6 eV compared to $\sim 36 \mathrm{eV}$ for gaseous ionization) allows use of very thin detectors with reasonable signal

Slide from Markus Elsing


## schema of silicon microstrip sensor

- reverse bias: backplane set to positive voltage (< 500 V )
- a traversing charged particle ionizes silicon, creating conduction electrons and holes that induce a measurable current by drifting to electrodes
- metal-semiconductor transition forms charge (Schottky) barrier similar to $p-n$ junction. Highly doped $n^{+}$layer reduces width of potential barrier and hence resistance


Tens of micron resolution from single hits! The alignment for all the modules gets quite ... tricky
https://atlas.web.cern.ch/Atlas/GROUPS/PHYSICS/PLOTS/IDTR-2017-004/



## Let's look very carefully at these plots together

## Challenges for silicon systems at the LHC

Just like your digital camera - except more directional, ie pointing in different directions (why?)... and we have to read out data a lot faster!

Not to mention dealing with intense radiation environments. Why might this be a problem? How can we mitigate the problem?


## ATLAS TRT (Transition Radiation Tracker)



300,000 straws ( 4 mm in diameter) at larger track radius, ~35 measurements per track. Also helps with particle identification

## Drift Tubes in ATLAS: Inner Detector and Muon Spectrometer

- classical detection technique for charged particles based on gas ionization and drift time measurement

example: segment in muon drift tubes reconstruction from measured drift circles (left-right ambiguity)


TRT: Kapton tubes, $\quad \varnothing=4 \mathrm{~mm}$
MDT: Aluminium tubes, $\varnothing=30 \mathrm{~mm}$

Slide from Markus Elsing

## Combining Tracking with PID: the ATLAS TRT

- e/ $\pi$ separation via transition radiation: polymer (PP) fibers/foils interleaved with drift tubes

total: 370 k straws

$$
\text { barrel }(|\eta|<0.7) \text { : }
$$

36 r- $\phi$ measurements / track resolution $\sim 130 \mu \mathrm{~m} /$ straw
14 end-cap wheels ( $|\eta|<2.1$ ):
40 or less $z-\phi$ points



Slide from Markus Elsing ${ }_{92}$

## ATLAS-CONF-2011-128



Given ~30 hits per track, small difference in probability to measure hits with a higher threshold can give significant particle identification/ background rejection for electron ID

## $p \propto B \cdot R$

For precise measurement of p, want precise measurement of radius of curvature. 100 GeV object at ATLAS has $R=166 \mathrm{~m}$ ! A 1 TeV pion will have R close to 1 km ... need precision in silicon systems of order $\sim 10$ microns per measurement


Particle travels $\sim \mathrm{c} \tau$ before decaying in its own rest frame but in lab frame, $\sim \gamma c \tau$. Typically interested in objects with substantial Lorentz gamma factors, so muons, neutrons and charged pions and kaons travel and decay in the detector

When a relativistic charged particle passes through some material with atomic number Z and electron density n , it ionizes the atoms and thus loses energy. I is the mean excitation energy in the material $\sim 10 \mathrm{Z}[\mathrm{eV}]$

$$
-\frac{d E}{d x}=\frac{4 \pi}{m_{e} c^{2}} \cdot \frac{n z^{2}}{\beta^{2}} \cdot\left(\frac{e^{2}}{4 \pi \varepsilon_{0}}\right)^{2} \cdot\left[\ln \left(\frac{2 m_{e} c^{2} \beta^{2}}{I \cdot\left(1-\beta^{2}\right)}\right)-\beta^{2}\right]
$$

Not very intuitive :)


Fig. 27.1: Stopping power $(=\langle-d E / d x\rangle)$ for positive muons in copper as a function of $\beta \gamma=p / M c$ over nine orders of magnitude in momentum (12 orders of magnitude in kinetic energy). Solid curves indicate the total stopping power. Data below the break at $\beta \gamma \approx 0.1$ are taken from ICRU 49 [4], and data at higher energies are from Ref. 5. Vertical bands indicate boundaries between different approximations discussed in the text. The short dotted lines labeled " $\mu^{-}$" illustrate the "Barkas effect," the dependence of stopping power on projectile charge at very low energies [6].

## From PDG. Particles with

 $\beta \gamma \sim 3$ lose the least amount of energy as they travel, and are referred to as "minimum ionizing" particles
## Using dE/dx in the pixel system of ATLAS


https://atlas.web.cern.ch/Atlas/GROUPS/PHYSICS/PLOTS/PIX-2015-002/

## Using dE/dx in the TRT system of ATLAS


https://twiki.cern.ch/twiki/bin/view/AtlasPublic/TRTPublicResults


Bremsstrahlung


Literally "braking radiation" in German, bremsstrahlung is the main way that electrons of interest at ATLAS lose energy. In EM you might have seen that the amount of energy radiated away goes as $1 /$ mass $^{2}$, which is why one reason why ee accelerators need to be so big, and why muons don't lose much energy in the detector

## Radiation length for high-energy electrons/photons



Radiation length $X_{0}$ is the average distance over which the energy of an electron is reduced by 1/e (characteristic distance scale for material effects)
$X_{0}$ (iron) $=1.8 \mathrm{~cm}$
$X_{0}($ lead $)=0.6 \mathrm{~cm}$

Very similar for photons



Bremsstrahlung


After $1 \mathrm{X}_{0}$ we have roughly twice as many particles, with half the original energy After $2 \mathrm{X}_{0}$, we have $\sim 4 \mathrm{x}$ as many particles, with 1/4 the original energy After $n X_{0}$, we have $\sim 2^{n}$ as many particles, and the energy of each is $\sim 2^{-n}$ of the original value At some critical energy $\mathrm{E}_{\mathrm{c}}$, the Brem process stops

$$
<E>_{n}=\frac{E}{2^{n}}=E_{c}
$$

$$
\ln E-\ln 2^{n}=\ln E_{c}
$$

$$
\ln E-n \ln 2=\ln E_{c}
$$

$$
n \ln 2=\ln \left(E / E_{c}\right)
$$

$$
n=\frac{\ln \left(E / E_{c}\right)}{\ln 2}
$$

Here n is the number of radiation lengths where the maximum number of particles is observed, aka where Brem stops (for lead, 100 GeV electron gives $13 \mathrm{X}_{0}$, which is just a few cm )

Need some dense material to absorb energy, and some active material to notice this and produce an output signal. Can be the same material, often is not. ATLAS EM calorimeter is lead-liquid argon

A view inside the ATLAS LAr calorimeter


Trickier, as hadrons can be electrically neutral but even then the constituents (quarks) have electric charge. And many hadrons have an overall electric charge

Defined by the nuclear interaction length $\left(\lambda_{I}\right)$, which is the mean distance between hadronic interactions.
$\lambda_{\mathrm{I}}($ lead $)=17 \mathrm{~cm}$ (compare with $X_{0}($ lead $\left.)=0.6 \mathrm{~cm}\right)$

Also tricky because neutral pions decay 99\% of the time to a pair of photons

## ATLAS tile calorimeter

Intersperse steel absorber with plastic scintillator tiles, plastic doped with organic material. When charged secondary particles emerge from the steel, they excite the doped material, and emit UV light that can be re-emitted as one color by a dye

<br>

## The ATLAS tile calorimeter

 The ATLAS tile calorimeter

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    ,


## What emerges from the calorimeters?

Neutrinos... we
"detect" them conservation in the plane perpendicular to the beam

And muons! Have a set of toroidal magnets (giving ATLAS its shape and name) and more muon systems on
outside of detector muon systems on
outside of detector

# Highly non-trivial B field. Need to monitor not just alignment but also field itself  (1) 




## er

 $+$

## Putting it all together




## Photon can appear as isolated cluster in the EM calorimeter with nothing in hadronic calorimeter behind it and nothing else nearby. NO charged track pointing at calorimeter. Careful: lots of neutral pions!




Need to be careful because almost half of the photons at ATLAS interact with material in the detector before reaching the calorimeter! These photons look similar, except there are one or two clusters of energy with charged tracks not coming from the original
interaction. How can there be only one track? Other one can be very soft (low momentum)

## ATLAS LEXPERIMENT

Run Number: 191190, Event Number: 19448322 Date: 2011-10-16 16:11:14 CEST
$\mathrm{m}_{\mathrm{YY}}=125.8 \mathrm{GeV}$


## Putting it all together



## Electron appears as

 isolated cluster in the EM calorimeter with nothing in hadronic calorimeter behind it and a charged track. Photons and charged hadrons look like this too!
## Electron ID at ATLAS


"Isolation" of nearby activity (in calorimeters and also tracking system) can help us to distinguish real electrons from fake electrons


Muons appear as isolated, minimally ionizing particles in the calorimeter, with charged tracks in both the inner detector and also the muon spectrometer. Typically not so many fake muons, but can have non-prompt muons from hadron decays

## Muon ID at ATLAS



## Beautiful invariant mass plot with early Run 2 ATLAS data

## Muon isolation




# Let's look at these two isolation plots (from muons inside the Z boson mass window) 

arXiv:1603.05598

## Putting it all together



Neutrinos are not measured directly but inferred by applying the conservation of momentum to all measured objects

## Aside about neutrinos

Are neutrinos Majorana


Any guesses?!
particles (ie their own antiparticles)? Different than Dirac particles (not their own anti-particles). Would violate electron flavor number! How to observe this?

## Using "MET" at ATLAS




## Using momentum imbalance to look for dark matter production

Note the nomenclature of "Missing energy" as part of "MET" even though energy is a scalar. Physicists are lazy!


Protons leave charged tracks, neutrons do not. Each deposit energy in both the
electromagnetic and hadronic calorimeters.
Both typically not produced on their own but with lots of other particles nearby inside of a jet


Typical jet has ~60\% of energy in charged particles (mostly pions), $\sim 30 \%$ in photons from neutral pion decay, and 10\% neutrals. On average, of course

In a process like $p p \rightarrow q q b a r$, quarks are flying apart in opposite directions. They do not form a bound state together, but the energy pulling them apart leads to radiation of gluons and lots of other quarks, which decay to other objects with QCD color charge, which decay, etc into a collection of particles called a jet

## Taus can decay to


an electron (18\%);
a muon (17\%);
a single charged pion and extra neutral pions decaying to photons (48\%);
three charged pions plus extra neutral pions decaying to photons (15\%)

Taus are often
reconstructed as a narrow collimated jet of 1 or 3 tracks
(Ignoring neutrinos)

## Using taus at ATLAS



# In many models of 

 Beyond Standard Model physics, new Higgs-like particles prefer to decay to heavy objects. Can have such objects decay to a pair of taus


Photon as a particle (classical). But of course, what we really mean is that the field is quantized (here providing an attractive force), and the quantized unit of the field transmits some momentum from one object to another

Photon are still exchanged here, but now they

## proton

 transmit a repulsive force. It's a bit counterintuitive to think of forces being transmitted (or even better, 'mediated') by particles, especially on a macroscopic scale, but we are anyway talking about single particles

Think of time "passing" to the right, so initially (the "initial state") we start with an electron. And at the end (in the "final state"), we also have an electron.

At some intermediate time, there is a photon interacting with this electron (whether it emits it or absorbs it). We can combine multiple such diagrams

Think of time "passing" to the right, so initially (the "initial state") we start with an electron. And at the end (in the "final state"), we also have an electron.

At some intermediate time,
Time $\longrightarrow$ there is a photon interacting with this electron (whether it emits it or absorbs it). We can combine multiple such diagrams


Time $\longrightarrow$

At some initial time, there were two separate electrons. At a later time, there were also two electrons. At an intermediate time, they were exchanging a photon between them (repulsive force).

Møller scattering


What happens if we just change the direction of time? Everyone turn your heads now:)

Does this remind you of anything from earlier in the course?


Time $\longrightarrow$

Electron vs positron should now be more obvious/automatic

Here I put things back such that time moves to the right. Note that in the initial time (on left) there is a fermion line moving backwards! This is an antiparticle moving forwards in time. And on the right there is a fermion line also going backwards, this is again an anti-particle

## Feynman diagrams

## Bhaba scattering



Note the second diagram below... start with $\mathrm{e}^{+} \mathrm{e}^{-}$ and end with $\mathrm{e}^{+} \mathrm{e}^{-} \ldots$ these two diagrams are both needed (same initial and same final states, so they must interfere quantum mechanically!)

Pair annihilation

$\mathrm{e}^{-} \mathrm{e}^{+} \rightarrow \mathrm{YY}$

Pair production
Compton scattering

$Y Y \rightarrow e^{-} e^{+}$

$\mathrm{e}^{-} \gamma \rightarrow \mathrm{Ye}^{-}$

Electrons here could be replaced by other objects with electric charge. Note that Feynman diagrams so far just look like pictorial ways to represent processes, but they can be used for much more, as we'll see


## Modification to Bhaba scattering (same initial and final states, different diagram)

## Let's take a closer look at these diagrams



Let's slice this up into different time bins Can cleanly divide into internal
and external particles

To: Single initial muon
$\mathrm{T}_{1}$ : Muon after interaction with photon, and that photon
$\mathrm{T}_{2}$ : Some external photon, muon after interaction with one or two photons, the internal photon
$\mathrm{T}_{3}$ : Muon after interaction with external photon, before reabsorbing the internal photon
$\mathrm{T}_{4}$ : Single final muon

## Let's take a closer look at these diagrams

There are eight different particles in this diagram. Four of them are external to the diagram and can be observed.

The other four are internal to the digram and are called virtual particles and do not have to have onshell mass

Bhaba scattering again


Virtual particles here (the photon in each diagram) cannot be observed, since the initial and final states are the same, so these much have interfering matrix elements


Have extended one line and shrunk another. Could have spaced things further apart or closer together - we do not differentiate between these sorts of things, nor the angle of the lines (we are not artists, here, or at least, I am not)


Again, can't distinguish these diagrams. So the observable scattering amplitude must be due to the sum of all such diagrams ... but there are an infinite number of them! Let's draw another few

$\mathrm{e}^{+} \mathrm{e}^{-} \rightarrow \mathrm{e}^{+} \mathrm{e}^{-} \rightarrow \mathrm{e}^{+} \mathrm{e}^{-}$

$\mathrm{e}^{+} \mathrm{e}^{-} \rightarrow \mathrm{e}^{+} \mathrm{e}^{-} \rightarrow \mathrm{e}^{+} \mathrm{e}^{-} \rightarrow \mathrm{e}^{+} \mathrm{e}^{-}$
$\mathrm{e}^{+} \mathrm{e}^{-} \rightarrow \mathrm{e}^{+} \mathrm{e}^{-}$in both cases, since intermediate states are not observable. You might ask why we're bothering to draw such pictures, but they will shortly become an invaluable computation tool. Each of these represents a single number!


You may have noticed that I started putting circles at each interaction point these are vertices and describe a point in space-time where an interaction occurs.


There is no such
photon-photon-photon vertex in the SM (of course, you should know that already - the photon does not carry electric charge)

Muons can emit multiple photons, but not at the same vertex (note that the solid lines here represent any particle with electric charge)

## Neutral pion decay

Pair annihilation again, with quarks


Now we can see why neutral pions decay 99\% of the time to a pair of photons, and also why this happens so quickly (any good answers?)

Start with common building block:


Note the similarity with the common building block of QED:


Gluon exchange between quark and anti-quark:


A few
fundamental differences, though, between QED and QCD...

Gluon exchange between quarks:


Remember that quarks have color (and so do gluons!) Try and follow the color lines (always a good thing to check)

Remember that gluons themselves have QCD color, so they self-couple!


Can have threegluon vertices


And fourgluon vertices


Each vertex describes an emission/absorption of a force carrier.

In QED, the probability of this is governed by the unit-less number $\alpha=1 / 137$, which is good, since it means the probability for many of these is small


In QCD, the probability of this is governed by the unit-less number $\alpha_{s}>1$, which means the diagrams with more
vertices contribute ... more!

## Running of the coupling constant



The value of the coupling constant is indeed strong, but only at large distances. At small distances (such as inside of a proton), the strong coupling constant gets very small and we can calculate quantities

The property that the coupling gets weaker at smaller distances (aka at larger energies) is known as asymptotic freedom

Nobel prize for Wilczek, Gross and Politzer


Fig. 2.1 Screening of a charge $q$ by a dielectric medium.

## In classical $\mathrm{E} \& \mathrm{M}$, the coupling is given by the charge q , and this is reduced at large distances by a dielectric

Bhaba scattering again


## Virtual diagrams contributing



## Running of the E\&M fine structure constant

## ~1/137 at zero momentum transfer

$\sim 1 / 127$ at the $Z$ boson mass!

These are vacuum polarization effects

## Running of the QCD strong coupling constant

## Can measure this in a variety of ways

## Use trijet mass to determine the strong coupling constant!





## Can measure this in a variety of ways

## Use ttbar production cross section instead!


arXiv: 1307.1907



## Masses of objects are also not single numbers!

## The "mass" of the

 top quark depends on the energy scale of the collision! Use the differential cross section vs $\mathrm{m}_{\mathrm{tt}}$ as a proxy for this

Measurement of strong coupling from V cross sections
CMS
$38 \mathrm{pb}^{-1}(7 \mathrm{TeV})+18.2 \mathrm{pb}^{-1}(8 \mathrm{TeV})$
--CT14

- HERAPDF2.0
$\star$ MMHT14
* NNPDF3.0
$\qquad$



Effects from quark polarization and gluon polarization give opposite effects!
$\mathrm{a}=2 \mathrm{f}-11 \mathrm{n}$ ( $\mathrm{f}=$ number of flavors, $\mathrm{n}=$ number colors) defines whether coupling increases or decreases at short distances
$\mathrm{f}=6, \mathrm{n}=3 \rightarrow \mathrm{a}<0 \rightarrow$ asymptotic freedom (upper limit on no more than 17 generations of quarks)

# where f here is any fermion ( $Z$ boson couples to all quarks, charged leptons and neutral leptons) 




Griffiths uses a jagged straight line for EW bosons, but I will not. Just be careful, regardless of your choice!
$\mathrm{e}^{-} \nu_{\mathrm{e}} \rightarrow \mathrm{e}^{-} \nu_{\mathrm{e}}$

t channel diagrams


We of course don't scatter the bare quarks, but quarks inside hadrons (here inside a neutron)


Remember that W bosons are the only ones involved in change of flavor or type of matter


Remember that W bosons are the only ones involved in change of flavor or type of matter. Example, a tau converts into a tau neutrino and a W boson:


Top quarks decay nearly $\sim 100 \%$ of the time to a W boson and a bottom quark. The W boson can decay to a quark-antiquark pair, or to a charged lepton and a neutrino

Check lepton flavor numbers
and EM
charge!


This is how a charged pion can decay to a muon (and an anti-muon neutrino). Could have guessed W boson involvement, because it involves a change of flavor (here charged pion decay is much slower than neutral pion decay)

Check lepton
flavor numbers
and EM
charge!


We of course don't scatter the bare quarks, but quarks inside hadrons (here inside a neutron) This is how neutrons betadecay to protons
n


How do we account for strangenesschanging weak interactions? (We know they have to come via the weak force, but so far we have only seen decays within a matter generation)
$\wedge \rightarrow \mathrm{p} \pi^{-}$
 of quarks that we observe, but to a slightly skewed set instead:

What we observe
What weak force couples to

$$
\binom{u}{d^{\prime}}\binom{c}{s^{\prime}}\binom{t}{b^{\prime}}
$$

$$
\left(\begin{array}{c}
d^{\prime} \\
s^{\prime} \\
b^{\prime}
\end{array}\right)=\left(\begin{array}{ccc}
V_{u d} & V_{u s} & V_{u b} \\
V_{c d} & V_{c s} & V_{c b} \\
V_{t d} & V_{t s} & V_{t b}
\end{array}\right)\left(\begin{array}{c}
d \\
s \\
b
\end{array}\right)
$$

The Kobayashi-Maskawa matrix
What we observe

## What weak force

 couples to$$
\binom{u}{d}\binom{c}{s}\binom{t}{b}
$$

$$
\binom{u}{d^{\prime}}\binom{c}{s^{\prime}}\binom{t}{b^{\prime}}
$$

$$
\left(\begin{array}{c}
d^{\prime} \\
s^{\prime} \\
b^{\prime}
\end{array}\right)=\left(\begin{array}{ccc}
V_{u d} & V_{u s} & V_{u b} \\
V_{c d} & V_{c s} & V_{c b} \\
V_{t d} & V_{t s} & V_{t b}
\end{array}\right)\left(\begin{array}{c}
d \\
s \\
b
\end{array}\right)
$$

\(\left($$
\begin{array}{c}d^{\prime} \\
s^{\prime} \\
b^{\prime}\end{array}
$$\right)=\left($$
\begin{array}{ccc}0.97434 & 0.22506 & 0.00357 \\
0.22492 & 0.97351 & 0.0411 \\
0.00875 & 0.0403 & 0.99915\end{array}
$$\right)\left(\begin{array}{l}d <br>
s <br>

b\end{array}\right)\)| 2017 |
| :--- |
| PDG |

Off-diagonal numbers are small, which has implications for which decays happen more/less frequently. The matrix has to be unitary. Note that l've cheated a bit, because there are phases in the matrix that I have ignored so far

$$
\left(\begin{array}{l}
d^{\prime} \\
s^{\prime} \\
b^{\prime}
\end{array}\right)=\left(\begin{array}{ccc}
0.97434 & 0.22506 & 0.00357 \\
0.22492 & 0.97351 & 0.0411 \\
0.00875 & 0.0403 & 0.99915
\end{array}\right)\left(\begin{array}{l}
d \\
s \\
b
\end{array}\right) \begin{aligned}
& 2017 \\
& \text { PDG }
\end{aligned}
$$

Boson couplings in weak theory




In general, can always replace a photon by a Z boson in a diagram. What about the reverse?

Boson couplings in weak theory





What are the s-channel and t-channel diagrams for $\mathrm{e}^{+} \mathrm{e}^{-} \rightarrow \mathrm{W}+\mathrm{W}$-? Which way do arrows go? Why? What about $\mathrm{e}^{+} \mathrm{e}^{-} \rightarrow \mu^{+} \mu^{-}$? What diagrams are there?

Particles... like to decay to lighter objects, unless there is a reason that they can't decay (such as a conservation law)

What is the most obvious reason for non-decay?

# $p^{+} \rightarrow n e^{+} \nu_{e}$ 

## Hint, look at the energy in the rest frame

Momentum conservation
Angular momentum conservation Conservation of charge Conservation of color

Conservation of baryon number
Conservation of lepton (electron/muon/tau) number
Quark flavor (not quite conserved)
https://www.ippp.dur.ac.uk/~krauss/Lectures/QuarksLeptons/QCD/Quarkonium 0.html


The phi decays much more commonly to K Kbar, despite a phase space preference for decays to three pions. Why? The OZI rule...
http://www.ippp.dur.ac.uk/~krauss/Lectures/QuarksLeptons/QCD/Equations/jpsidecay.gif


OZI rule explains why the J/Psi takes so long to decay. What does this have to do with asymptotic freedom?


If we can make that cut, it means all the energy of the quarks is temporarily in the form of gluons. But QCD couplings decrease at high energy, so there is less probability of this happening


## Explained by the OZI rule, let's go over this

## $\phi(1020)$ DECAY MODES

|  | Mode | Fraction $\left(\Gamma_{i} / \Gamma\right)$ | Scale factor/ <br> Confidence level |
| :--- | :--- | ---: | ---: | ---: |
| $\Gamma_{1}$ | $K^{+} K^{-}$ | $(48.9 \quad \pm 0.5) \%$ | $\mathrm{~S}=1.1$ |
| $\Gamma_{2}$ | $K_{L}^{0} K_{S}^{0}$ | $(34.2 \quad \pm 0.4) \%$ | $\mathrm{~S}=1.1$ |
| $\Gamma_{3}$ | $\rho \pi+\pi^{+} \pi^{-} \pi^{0}$ | $(15.32 \pm 0.32) \%$ | $\mathrm{~S}=1.1$ |
| $\Gamma_{4}$ | $\rho \pi$ |  |  |
| $\Gamma_{5}$ | $\pi^{+} \pi^{-} \pi^{0}$ | https://pdg.lbl.gov/2012/listings/rpp2012-list-phi-1020.pdf |  |

## Photons can't decay (they

 already have zero mass nothing lighter to decay into)Protons are the lightest baryon, so baryon conservation tells us that they are stable

This is why the world is made up of protons, neutrons (which are stable inside nucleii), electrons and neutrinos

Decays are probabilistic - you never know when an object will decay, but you can say on average how long a type of object will take to decay

```
\mu: 2.2\times10-6 s
\pi+: 2.6\times10-8 s
\pi}0:8.3\times1\mp@subsup{0}{}{-17}\textrm{s
\tau: 2.9\times10-13 s
n (free): }880\mathrm{ s
top quark: 4\times10-25 s
```

Many of these objects will decay to multiple final states with different, predicted branching ratios

Strong decays: $10^{-23} \mathrm{~s}$ (top quark)
EM decays: 10-16 s ( $\pi^{0}$ )
Weak decays: longer (neutron and muon)

The more (less) phase space for a decay, the faster (slower) it will occur. See the free neutron



## OK, not a real physics explanation, but a good way to remember things

## Griffiths problems 1.19, 2.1, 2.2, 2.5, 2.6, 2.7, <br> 2.8

