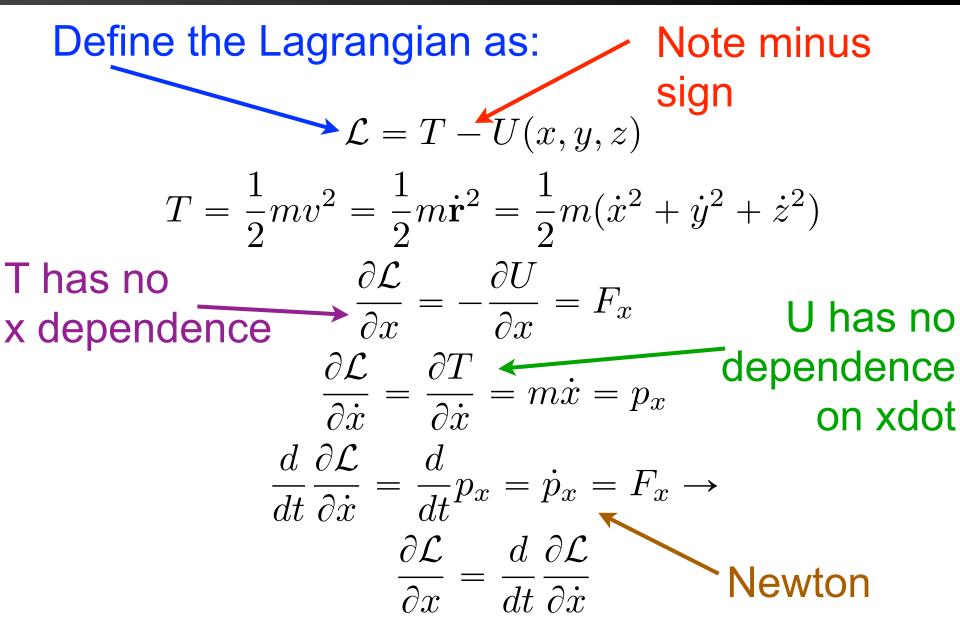
But what does this tell us about mechanics?



$\partial \mathcal{L}$	$d \partial \mathcal{L}$
$\overline{\partial x} =$	$\overline{dt} \ \overline{\partial \dot{x}}$
$\partial \mathcal{L}$	$d \partial \mathcal{L}$
$\overline{\partial y} =$	$\overline{dt} \ \overline{\partial \dot{y}}$
$\partial \mathcal{L}$	$d \partial \mathcal{L}$
$\overline{\partial z} =$	$\overline{dt} \ \overline{\partial \dot{z}}$

From last slide. Repeat to get the other two equations. These are in form of Euler-Lagrange equations!

Hamilton's principle:

 $S = \text{Action integral} = \int_{t_1}^{t_2} \mathcal{L} dt = \int_{t_1}^{t_2} (T - U) dt$ is stationary

Keys to using Lagrange's Equations

$\partial \mathcal{L} ~~~ d~~ \partial \mathcal{L}$	Can use more	$\partial \mathcal{L} ~~~ d~~ \partial \mathcal{L}$
$\overline{\partial x} = \overline{dt} \overline{\partial \dot{x}}$	generalized	$\overline{\partial q_1} = \overline{dt} \overline{\partial \dot{q_1}}$
$\partial \mathcal{L} _ d \partial \mathcal{L}$	coordinates	$\partial \mathcal{L} \ _ \ d \ \partial \mathcal{L}$
$\overline{\partial y} = \overline{dt} \ \overline{\partial \dot{y}}$		$\overline{\partial q_2} = \overline{dt} \overline{\partial \dot{q_2}}$
$\partial \mathcal{L}$ $d \ \partial \mathcal{L}$		$\frac{\partial \mathcal{L}}{\partial \mathcal{L}} = \frac{d}{\partial \mathcal{L}} \frac{\partial \mathcal{L}}{\partial \mathcal{L}}$
$\overline{\partial z} = \overline{dt} \overline{\partial \dot{z}}$		$\partial q_3 \stackrel{-}{} dt \ \partial \dot{q_3}$

q_i (generalized coordinates)
 can really be any coordinates
 in any frame, but remember to Why is this?
 write down
 Lagrangian in an inertial frame

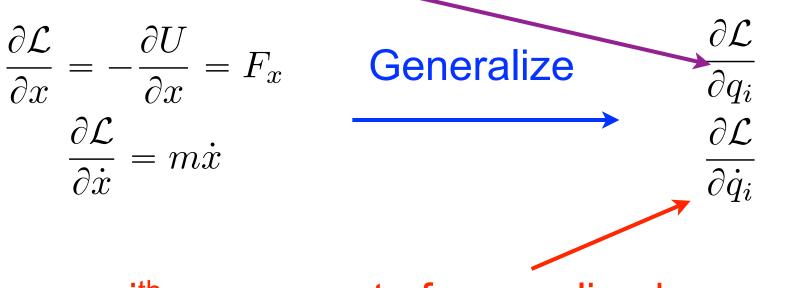
Keys to using Lagrange's Equations

$\partial \mathcal{L} \qquad d \ \partial \mathcal{L}$	$\partial \mathcal{L} ~~~~~~~~~~~~~~~~~~~~~~~~~~~~~~~~~~~~$
$\overline{\partial q_1} = \overline{dt} \overline{\partial \dot{q_1}}$	$\overline{\partial p_1} = \overline{dt} \overline{\partial \dot{p_1}}$
$\partial \mathcal{L} \qquad d \ \partial \mathcal{L}$	$\partial \mathcal{L} ~~~ d~~ \partial \mathcal{L}$
$\overline{\partial q_2} = \overline{dt} \overline{\partial \dot{q_2}}$	$\frac{\partial p_2}{\partial p_2} = \frac{\partial t}{\partial t} \frac{\partial \dot{p_2}}{\partial \dot{p_2}}$
$\partial \mathcal{L} \qquad d \ \partial \mathcal{L}$	$\partial \mathcal{L} = d \ \partial \mathcal{L}$
$\overline{\partial q_3} = \overline{dt} \overline{\partial \dot{q_3}}$	$\frac{\partial p_3}{\partial p_3} = \frac{\partial p_3}{\partial t} \frac{\partial p_3}{\partial p_3}$

Additional generalized coordinates for every new particle. Can use, for example, center of mass position and relative position from CM A particle moves in a conservative force field in two dimensions. What are the equations of motion, using Cartesian coordinates?

$$\mathcal{L} = \mathcal{L}(x, y, \dot{x}, \dot{y}) = T - U = \frac{1}{2}m(\dot{x}^2 + \dot{y}^2) - U(x, y)$$
$$\frac{\partial \mathcal{L}}{\partial x} = \frac{d}{dt}\frac{\partial \mathcal{L}}{\partial \dot{x}}$$
$$\frac{\partial \mathcal{L}}{\partial x} = -\frac{\partial U}{\partial x} = F_x$$
$$\frac{\partial \mathcal{L}}{\partial \dot{x}} = m\dot{x}$$
$$\frac{\partial \mathcal{L}}{\partial \dot{x}} = \frac{d}{dt}\frac{\partial L}{\partial \dot{x}} \rightarrow F_x = m\ddot{x}$$
similarly $F_y = m\ddot{y}$

ith component of generalized force



ith component of generalized momentum

 $\frac{\partial \mathcal{L}}{\partial q_i} = \frac{d}{dt} \frac{\partial \mathcal{L}}{\partial \dot{q_i}}$

And Lagrange says the generalized force = rate of change of generalized momentum Another examples (following Taylor)

A particle moves in a conservative force field in two dimensions. What are the equations of motion, using Polar coordinates?

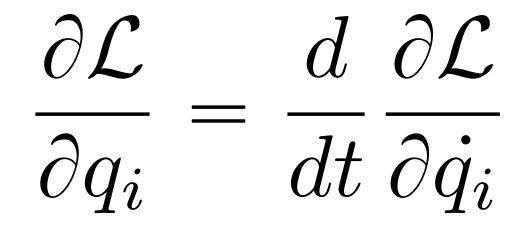
$$\mathcal{L} = \mathcal{L}(r, \phi, \dot{r}, \dot{\phi}) = T - U = \frac{1}{2}m(\dot{r}^2 + r^2\dot{\phi}^2) - U(r, \phi)$$
$$\frac{\partial \mathcal{L}}{\partial r} = \frac{d}{dt}\frac{\partial \mathcal{L}}{\partial \dot{r}}$$
$$mr\dot{\phi}^2 - \frac{\partial U}{\partial r} = \frac{d}{dt}m\dot{r} = m\ddot{r}$$
$$mr\dot{\phi}^2 + F_r = \frac{d}{dt}m\dot{r} = m\ddot{r}$$
$$F_r = m(\ddot{r} - r\dot{\phi}^2)$$
$$\mathbf{F}_r = \mathbf{mar}$$

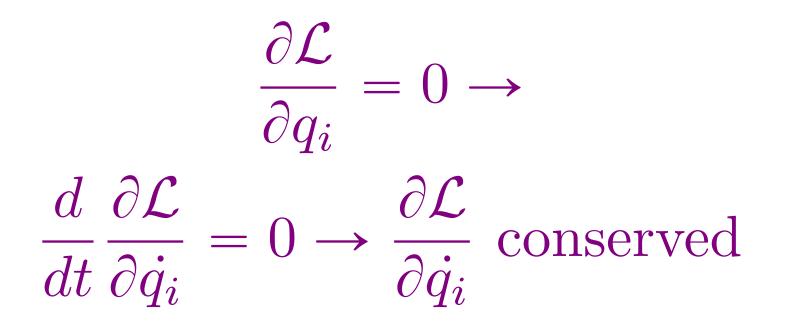
Continuing on

 $\frac{\partial \mathcal{L}}{\partial \phi} = \frac{d}{dt} \frac{\partial \mathcal{L}}{\partial \dot{\phi}}$ $I\omega = L$ $-\frac{\partial U}{\partial \phi} = \frac{d}{dt} (mr^2 \dot{\phi})$ $\mathbf{F} = \nabla U$ $F_{\phi} = (\nabla U)_{\phi}$ $\nabla U = \frac{\partial U}{\partial r} \hat{\mathbf{r}} + \frac{1}{r} \frac{\partial U}{\partial \phi} \hat{\phi}$ $(\nabla U)_{\phi} = \frac{1}{r} \frac{\partial U}{\partial \phi}$ $\frac{\partial U}{\partial \phi} = r(\nabla U)_{\phi}$

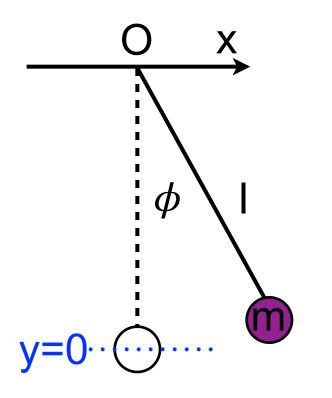
 $\frac{\partial U}{\partial \phi} = r(\nabla U)_{\phi} = -rF_{\phi}$ $rF_{\phi} = \frac{d}{dt}I\omega = \frac{d}{dt}L$ Torque = $\Gamma = \frac{dL}{dt}$

A useful thing to have spotted





And work on problem 7.1 yourself, and then we'll work on 7.2 and 7.17 together



Two dimensional system, but really need only one coordinate to describe the system due to constraints of the rod (since rod length is constant)

More examples from Taylor (chapter 7.2)

$$\mathcal{L} = T - U = \frac{1}{2}m(l^{2}\dot{\phi}^{2}) - mgy$$

$$\mathcal{L} = T - U = \frac{1}{2}m(l^{2}\dot{\phi}^{2}) - mgl(1 - \cos\phi)$$

$$\frac{\partial \mathcal{L}}{\partial \phi} = \frac{d}{dt}\frac{\partial \mathcal{L}}{\partial \dot{\phi}}$$

$$-mgl\sin\phi = \frac{d}{dt}(ml^{2}\dot{\phi}) = ml^{2}\ddot{\phi}$$

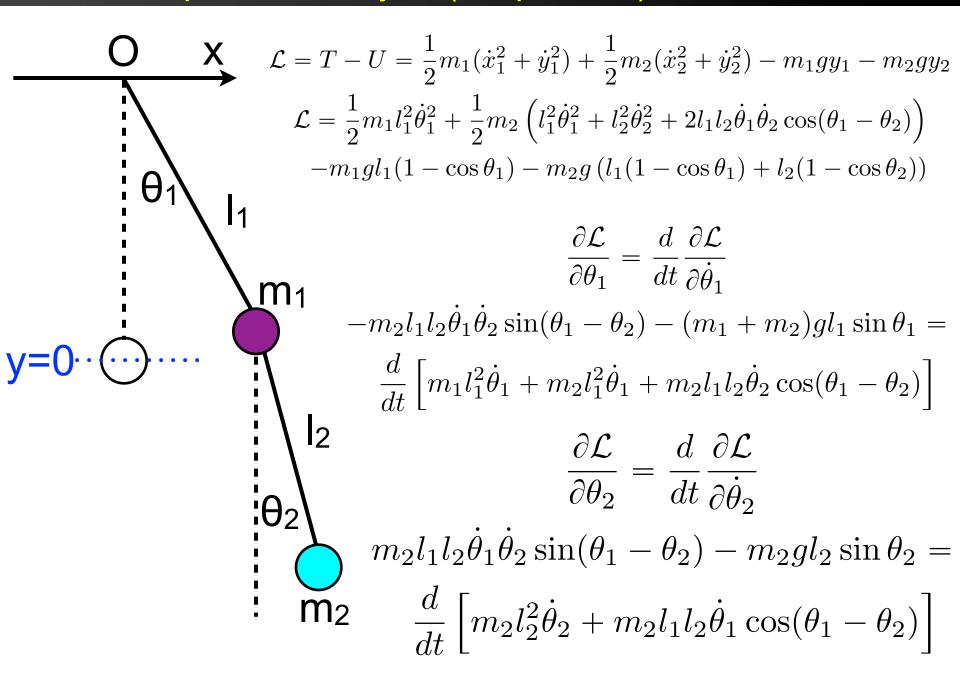
$$\int_{\mathbf{T} \text{orque on mass about O}} \int_{\mathbf{\phi}} ml^{2} = I$$

$$\ddot{\phi} = \alpha$$

$$\Gamma = I\alpha$$

More examples from Taylor (chapter 7.3), just to see...

More examples from Taylor (chapter 7.3)



More examples from Taylor (chapter 7.3)

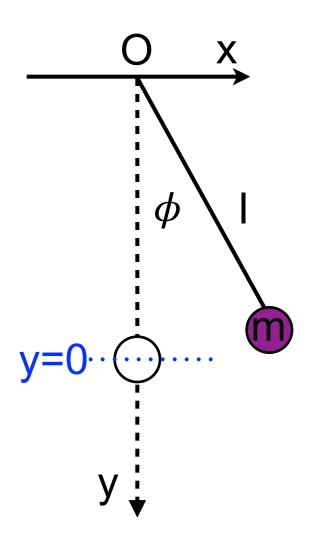
No full solutions, but we'll get to Χ small-angle solutions in Chapter 11. Imagine doing this with Newton, though? $\frac{\partial \mathcal{L}}{\partial \theta_1} = \frac{d}{dt} \frac{\partial \mathcal{L}}{\partial \dot{\theta}_1}$ m_1 $-m_2 l_1 l_2 \dot{\theta}_1 \dot{\theta}_2 \sin(\theta_1 - \theta_2) - (m_1 + m_2) g l_1 \sin \theta_1 =$ **y=0**··(···) $\frac{d}{dt} \left[m_1 l_1^2 \dot{\theta}_1 + m_2 l_1^2 \dot{\theta}_1 + m_2 l_1 l_2 \dot{\theta}_2 \cos(\theta_1 - \theta_2) \right]$ \mathbf{I}_2 $\frac{\partial \mathcal{L}}{\partial \theta_2} = \frac{d}{dt} \frac{\partial \mathcal{L}}{\partial \dot{\theta}_2}$ θ_2 $m_2 l_1 l_2 \dot{\theta}_1 \dot{\theta}_2 \sin(\theta_1 - \theta_2) - m_2 g l_2 \sin \theta_2 =$ $\widetilde{\mathbf{m}}_{2} \qquad \frac{d}{dt} \left[m_{2}l_{2}^{2}\dot{\theta}_{2} + m_{2}l_{1}l_{2}\dot{\theta}_{1}\cos(\theta_{1}-\theta_{2}) \right]$

$$\frac{\partial \mathcal{L}}{\partial q_i} = \frac{d}{dt} \frac{\partial \mathcal{L}}{\partial \dot{q_i}} \quad \text{Let's say we chose} \\ \mathbf{q_1} = \mathbf{\theta_{1, q_2}} = \mathbf{\theta_2}$$

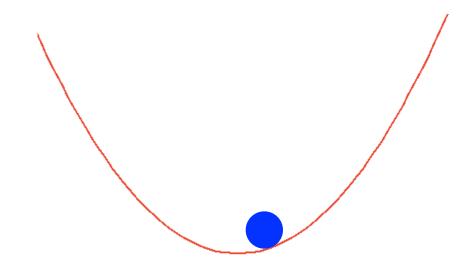
$$\mathbf{r}_1 = x_1 \hat{\mathbf{x}_1} + y_1 \hat{\mathbf{y}_1}$$
$$\mathbf{r}_2 = x_2 \hat{\mathbf{x}_2} + y_2 \hat{\mathbf{y}_2}$$
with

 $x_1 = l_1 \sin \theta_1$ $y_1 = l_1 \cos \theta_1$ $x_2 = l_1 \sin \theta_1 + l_2 \sin \theta_2$ $y_2 = l_1 \cos \theta_1 + l_2 \cos \theta_2$

These relationships do not depend on time, so they are referred to as being natural

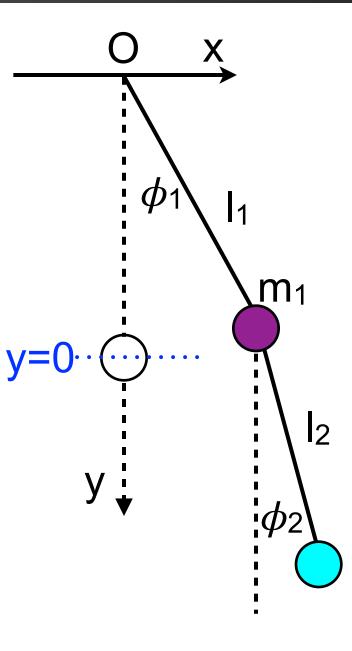


Two dimensional system, but really need only one coordinate to describe the system due to constraints of the rod. We call such systems constrained

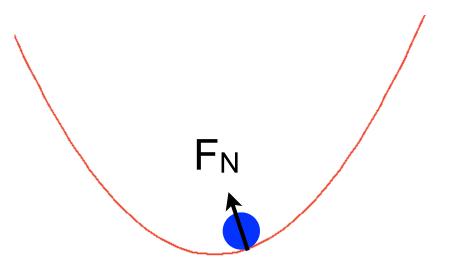


This system, with a ball moving along this bead, is also constrained

And yet a bit more terminology

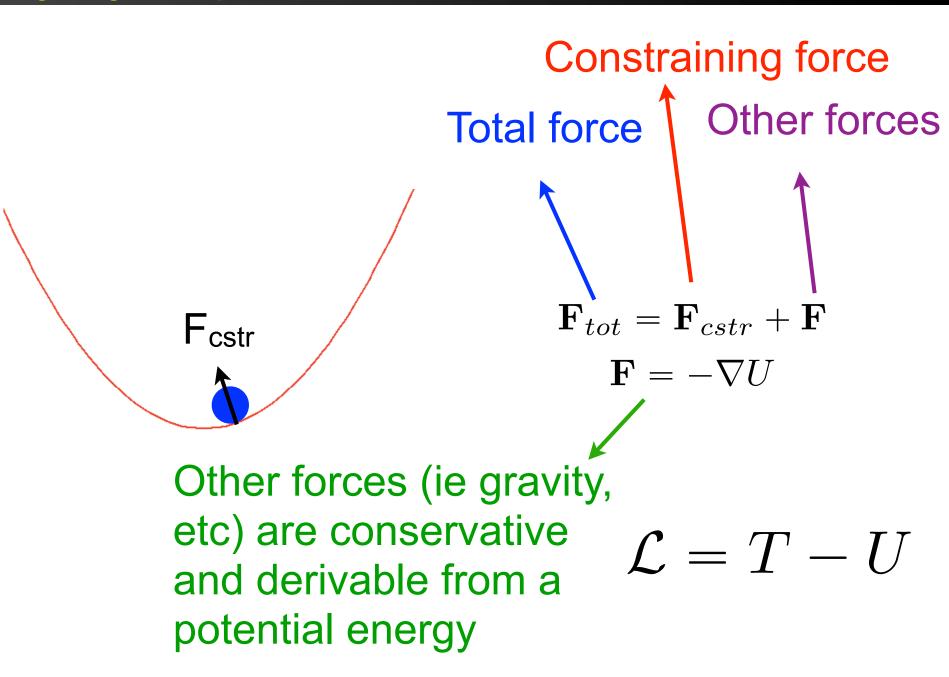


System has two degrees of freedom, and is described by two generalized coordinates. If these two numbers are equal, the system is holonomic



Ball has (for example) gravity pulling it down, but also a constraint force/normal force that keeps it moving along the parabola. Why doesn't this matter?

Lagrange's Equations with constraints



Recall Hamilton's principle:

 $S = \text{Action integral} = \int_{t_1}^{t_2} \mathcal{L}dt = \int_{t_1}^{t_2} (T - U)dt$ is stationary

Particle passes through \mathbf{r}_1 and \mathbf{r}_2 at t_1 and t_2

Another
$$\mathbf{r}(t) = \mathbf{r}(t) + \epsilon(t)$$

(potentially vrong) path Right path

$$\epsilon(t_1) = \epsilon(t_2) = 0$$

Define action for right/wrong paths

$$S_{0} = \int_{t_{1}}^{t_{2}} T(\mathbf{r}, \dot{\mathbf{r}}, t) - U(\mathbf{r}, \dot{\mathbf{r}}, t) dt \longrightarrow \begin{array}{l} \text{action for the correct path} \\ S_{0} = \int_{t_{1}}^{t_{2}} \frac{m}{2} \dot{\mathbf{r}}^{2} - U(\mathbf{r}, \dot{\mathbf{r}}, t) dt \\ S = \int_{t_{1}}^{t_{2}} T(\mathbf{R}, \dot{\mathbf{R}}, t) - U(\mathbf{R}, \dot{\mathbf{R}}, t) dt \longrightarrow \begin{array}{l} \text{action for arbitrary path} \\ S = \int_{t_{1}}^{t_{2}} \frac{m}{2} \dot{\mathbf{R}}^{2} - U(\mathbf{R}, \dot{\mathbf{R}}, t) dt \\ S = \int_{t_{1}}^{t_{2}} \frac{m}{2} (\dot{\mathbf{r}} + \dot{\epsilon})^{2} - U(\mathbf{R}, \dot{\mathbf{R}}, t) dt \\ S - S_{0} = \delta S = \int_{t_{1}}^{t_{2}} \frac{m}{2} \dot{\epsilon}^{2} + m \dot{\mathbf{r}} \cdot \dot{\epsilon} - \left(U(\mathbf{R}, \dot{\mathbf{R}}, t) - U(\mathbf{r}, \dot{\mathbf{r}}, t) \right) dt \\ \end{array}$$

Follow similar strategy as before

 δ

And once again integrate by parts and use boundary conditions

$$\delta S = \int_{t_1}^{t_2} [m\dot{\mathbf{r}} \cdot \dot{\epsilon} - \nabla U(\mathbf{r}, \dot{\mathbf{r}}, t) \cdot \epsilon] dt$$
Integrate
by parts:
$$\int_{a}^{b} u dv = [uv]_{a}^{b} - \int v du$$

$$u = m\dot{\mathbf{r}}, dv = \dot{\epsilon} dt$$

$$\int_{t_1}^{t_2} (m\dot{\mathbf{r}} \cdot \dot{\epsilon}) dt = [m\dot{\mathbf{r}} \cdot \epsilon]_{t_1}^{t_2} - \int_{t_1}^{t_2} (m\epsilon \cdot \ddot{\mathbf{r}}) dt$$
Zero because
$$\varepsilon(\mathbf{t}_1) = \varepsilon(\mathbf{t}_2) = 0$$

$$\delta S = \int_{a}^{t_2} [-m\epsilon \cdot \ddot{\mathbf{r}} - \nabla U \cdot \epsilon] dt$$

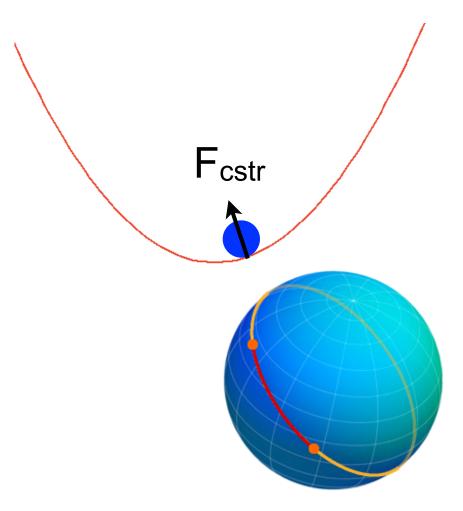
 J_{t_1}

Working out the math

$$\delta S = \int_{t_1}^{t_2} \left[-m\epsilon \cdot \ddot{\mathbf{r}} - \nabla U \cdot \epsilon \right] dt$$
$$\delta S = -\int_{t_1}^{t_2} \epsilon \cdot \left[m\ddot{\mathbf{r}} + \nabla U \right] dt$$
$$m\ddot{\mathbf{r}} = \mathbf{F}_{tot} = \mathbf{F}_{cstr} + \mathbf{F} = \mathbf{F}_{cstr} - \nabla U$$
$$\delta S = -\int_{t_1}^{t_2} \epsilon \cdot \left[\mathbf{F}_{cstr} - \nabla U + \nabla U \right] dt$$
$$\delta S = -\int_{t_1}^{t_2} (\epsilon \cdot \mathbf{F}_{cstr}) dt$$

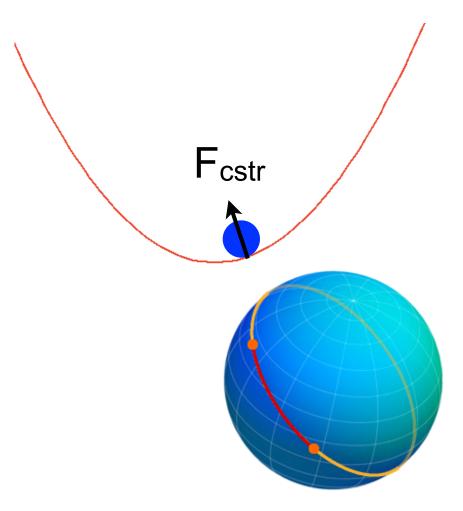
So we are left thinking about the dot product between ϵ and the constraining force

So we are left thinking about the dot product between ϵ and the constraining force



 ϵ is a variation on the path that the particle could take. BUT it can only take a path along the surface allowed by the constraint

So we are left thinking about the dot product between ϵ and the constraining force



The constraining force is a **normal force**, ie perpendicular to the volume/surface in which the particle can move. So ...

 $\epsilon \cdot \mathbf{F}_{cstr} = 0 \to \delta S = 0$

The action integral is stationary at the right path, which means that Euler-Lagrange equations still apply

Remember that this only works if the constraining force is a normal force to the surface of motion You should try and become as comfortable as possible working out problems in the Lagrangian formalism, so let's solve examples 7.4-7.7 on the board together, and then problems 7.16,7.20,7.23,7.34

Taylor 7.3, 7.8, 7.14, 7.18, 7.29